

The Higgs sector

Recall the scalar potential (at tree-level):

$$\begin{aligned}
 V = & (|\mu|^2 + m_{H_u}^2) (|H_u^0|^2 + |H_u^+|^2) + (|\mu|^2 + m_{H_d}^2) (|H_d^0|^2 + |H_d^-|^2) \\
 & + [b (H_u^+ H_d^- - H_u^0 H_d^0) + \text{h.c.}] \\
 & + \frac{1}{8} (g^2 + g'^2) (|H_u^0|^2 + |H_u^+|^2 - |H_d^0|^2 - |H_d^-|^2)^2 \\
 & + \frac{1}{2} g^2 |H_u^+ H_d^{0*} + H_u^0 H_d^{-*}|^2
 \end{aligned} \tag{1}$$

Minimize \rightarrow physical Higgs masses and the h^0 - H^0 mixing angle α :

$$m_{A^0}^2 = 2b / \sin 2\beta \tag{2}$$

$$m_{H^\pm}^2 = m_{A^0}^2 + m_W^2 \tag{3}$$

$$m_{h^0, H^0}^2 = \frac{1}{2} \left(m_{A^0}^2 + m_Z^2 \mp \sqrt{(m_{A^0}^2 + m_Z^2)^2 - 4m_Z^2 m_{A^0}^2 \cos^2 2\beta} \right) \tag{4}$$

$$\frac{\sin 2\alpha}{\sin 2\beta} = -\frac{m_{A^0}^2 + m_Z^2}{m_{H^0}^2 - m_{h^0}^2} \quad \frac{\cos 2\alpha}{\cos 2\beta} = -\frac{m_{A^0}^2 - m_Z^2}{m_{H^0}^2 - m_{h^0}^2} \tag{5}$$

How many free parameters are there?

- b term
- $(|\mu|^2 + m_{H_u}^2)$
- $(|\mu|^2 + m_{H_d}^2)$

One combination determines $v_u^2 + v_d^2 = v^2$ – already known from the Z mass. That leaves two free parameter combinations:

usually chosen to be m_{A^0} and $\tan \beta$.

This only works at tree level. Once radiative corrections are included, other SUSY parameters enter into the Higgs sector. E.g.:

$$\Delta(m_{h^0}^2) \simeq \frac{3}{4\pi^2} v^2 y_t^4 \sin^4 \beta \ln \left(\frac{m_{\tilde{t}_1} m_{\tilde{t}_2}}{m_t^2} \right) \tag{6}$$

Higgs couplings to SM fermions

Recall we had

$$\mathcal{L} = -y_t \bar{u}_3 Q H_u + y_b \bar{d}_3 Q H_d + y_\tau \bar{e}_3 L H_d \quad (7)$$

$$= -y_t (\bar{t} t H_u^0 - \bar{t} b H_u^+) + y_b (\bar{b} t H_d^- - \bar{b} b H_d^0) + y_\tau (\bar{\tau} \nu_\tau H_d^- - \bar{\tau} \tau H_d^0) \quad (8)$$

Using $v_u = \langle H_u^0 \rangle v \sin \beta$ and $v_d = \langle H_d^0 \rangle = v \cos \beta$ and $m_W = gv/\sqrt{2}$, we can solve for the Yukawa couplings in terms of the fermion masses:

$$y_t = \frac{gm_t}{\sqrt{2}m_W \sin \beta} \quad y_b = \frac{gm_b}{\sqrt{2}m_W \cos \beta} \quad y_\tau = \frac{gm_\tau}{\sqrt{2}m_W \cos \beta} \quad (9)$$

If $\tan \beta \gg 1$ then y_b and y_τ get enhanced.

Couplings of the Higgs mass eigenstates:

$$g_{h^0 \bar{t} t} = \frac{gm_t \cos \alpha}{2m_W \sin \beta} \quad g_{h^0 \bar{b} b} = -\frac{gm_b \sin \alpha}{2m_W \cos \beta} \quad (10)$$

$$g_{H^0 \bar{t} t} = \frac{gm_t \sin \alpha}{2m_W \sin \beta} \quad g_{H^0 \bar{b} b} = \frac{gm_b \cos \alpha}{2m_W \cos \beta} \quad (11)$$

$$g_{A^0 \bar{t} t} = \frac{igm_t}{2m_W} \cot \beta \gamma^5 \quad g_{A^0 \bar{b} b} = \frac{igm_b}{2m_W} \tan \beta \gamma^5 \quad (12)$$

$$g_{H^\pm \bar{t}_R b_L} = \frac{gm_t}{\sqrt{2}m_W} \cot \beta \quad g_{H^\pm \bar{t}_L b_R} = \frac{gm_b}{\sqrt{2}m_W} \tan \beta \quad (13)$$

Decoupling limit

An interesting limit occurs for $m_{A^0} \gg m_Z$: the so-called **decoupling limit**.

In the decoupling limit:

- $m_{A^0} \simeq m_{H^0} \simeq m_{H^\pm} \gg m_Z$
- m_{h^0} saturates its upper bound; $m_{h^0} \simeq m_Z |\cos 2\beta|$ at tree level
- $\alpha \simeq \beta - \pi/2$:
 - A^0, H^0, H^\pm live together in one linear combination of H_u and H_d
 - h^0 lives in the other linear combination of H_u and H_d , which also contains the Goldstone bosons G^0, G^\pm and the vev $v = \sqrt{v_u^2 + v_d^2}$

Standard Model Higgs

Then the Higgs sector looks like:

- One light SM-like Higgs h^0
- A multiplet A^0, H^0, H^\pm of heavy Higgses that don't affect the low-energy physics very much.

Need high-precision measurements of the h^0 couplings to distinguish it from the SM!

The details:

$\cos(\beta - \alpha)$ goes to zero in the limit $m_{A^0} \gg m_Z$:

$$\cos(\beta - \alpha) \simeq \frac{1}{2} \sin 4\beta \frac{m_Z^2}{m_{A^0}^2} \quad (14)$$

The h^0 couplings can be rewritten in a useful form in terms of $\cos(\beta - \alpha)$:

$$g_{h^0 \bar{t}t} = \frac{gm_t}{2m_W} [\sin(\beta - \alpha) + \cot \beta \cos(\beta - \alpha)] \quad (15)$$

$$g_{h^0 \bar{b}b} = \frac{gm_b}{2m_W} [\sin(\beta - \alpha) - \tan \beta \cos(\beta - \alpha)] \quad (16)$$

$$g_{h^0 W^+W^-} = gm_W \sin(\beta - \alpha) \quad (17)$$

$$g_{h^0 ZZ} = \frac{gm_Z}{\cos \theta_W} \sin(\beta - \alpha) \quad (18)$$

These all approach their SM values in the limit $\cos(\beta - \alpha) \rightarrow 0$.

Neutralinos and Charginos

The Higgsinos and the electroweak gauginos mix with each other because of the effects of electroweak symmetry breaking.

- Neutral Higgsinos $\tilde{H}_u^0, \tilde{H}_d^0$ and neutral gauginos \tilde{B}, \tilde{W}^0 mix:
form four neutral mass eigenstates called **neutralinos**, $\tilde{N}_{1,2,3,4}$ or $\tilde{\chi}_{1,2,3,4}^0$.
 - **Convention:** $m_{\tilde{N}_1} < m_{\tilde{N}_2} < m_{\tilde{N}_3} < m_{\tilde{N}_4}$
 - The lightest neutralino \tilde{N}_1 is usually assumed to be the LSP.
- Charged Higgsinos $\tilde{H}_u^\pm, \tilde{H}_d^\pm$ and winos \tilde{W}^\pm mix:
form two charged mass eigenstates called **charginos**, $\tilde{C}_{1,2}$ or $\tilde{\chi}_{1,2}^\pm$.
 - **Convention:** $m_{\tilde{C}_1} < m_{\tilde{C}_2}$

Neutralino mass terms:

$$\mathcal{L} = -\frac{1}{2} (\psi^0)^T \mathbf{M}_{\tilde{N}} \psi^0 + \text{h.c.} \quad (19)$$

where $\psi^0 = (\tilde{B}, \tilde{W}^0, \tilde{H}_d^0, \tilde{H}_u^0)$.

The mass matrix is:

$$\mathbf{M}_{\tilde{N}} = \begin{pmatrix} M_1 & 0 & -c_\beta s_W m_Z & s_\beta s_W m_Z \\ 0 & M_2 & c_\beta c_W m_Z & -s_\beta c_W m_Z \\ -c_\beta s_W m_Z & c_\beta c_W m_Z & 0 & -\mu \\ s_\beta s_W m_Z & -s_\beta c_W m_Z & -\mu & 0 \end{pmatrix} \quad (20)$$

Abbreviations: $s_\beta = \sin \beta$, $c_\beta = \cos \beta$, $s_W = \sin \theta_W$, $c_W = \cos \theta_W$

Danger: Some people use a different sign convention for μ . Looking for the $-\mu$ in the neutralino mass matrix tells you which convention they're using.

The neutralino mass matrix is diagonalized by a unitary matrix \mathbf{N} : $\tilde{N}_i = \mathbf{N}_{ij}\psi_j^0$:

$$\mathbf{M}_{\tilde{N}}^{diag} = \mathbf{N}^* \mathbf{M}_{\tilde{N}} \mathbf{N}^{-1} \quad (21)$$

where $\mathbf{M}_{\tilde{N}}^{diag}$ is defined to have positive real entries along the diagonal:

$$\mathbf{M}_{\tilde{N}}^{diag} = \begin{pmatrix} m_{\tilde{N}_1} & & & \\ & m_{\tilde{N}_2} & & \\ & & m_{\tilde{N}_3} & \\ & & & m_{\tilde{N}_4} \end{pmatrix} \quad (22)$$

$m_{\tilde{N}_i}$ are the absolute values of the eigenvalues of $\mathbf{M}_{\tilde{N}}$

or equivalently, $m_{\tilde{N}_i}^2$ are the eigenvalues of $\mathbf{M}_{\tilde{N}}^\dagger \mathbf{M}_{\tilde{N}}$

Caution: some people use a convention in which $m_{\tilde{N}_i}$ can be negative or even complex, to get a simpler mixing matrix \mathbf{N} .

$\mathbf{M}_{\tilde{N}}$ has 4 free parameters: M_1 , M_2 , μ , and $\tan\beta$.

If gaugino mass unification holds, so that $M_1 \simeq 0.5M_2$ due to the RGEs, this reduces to 3.

If $m_Z \ll |\mu \pm M_1|, |\mu \pm M_2|$ then the gaugino-Higgsino mixing is small.

Then (assuming also $M_1 < M_2 < |\mu|$)

$$\tilde{N}_1 \approx \tilde{B}, \quad m_{\tilde{N}_1} \approx M_1 + \dots \quad (23)$$

$$\tilde{N}_2 \approx \tilde{W}^0, \quad m_{\tilde{N}_2} \approx M_2 + \dots \quad (24)$$

$$\tilde{N}_3, \tilde{N}_4 \approx \frac{1}{\sqrt{2}} \left(\tilde{H}_u^0 \pm \tilde{H}_d^0 \right), \quad m_{\tilde{N}_3}, m_{\tilde{N}_4} \approx |\mu| + \dots \quad (25)$$

Chargino mass terms:

$$\mathcal{L} = -\frac{1}{2} (\psi^\pm)^T \mathbf{M}_{\tilde{C}} \psi^\pm + \text{h.c.} \quad (26)$$

where $\psi^\pm = (\tilde{W}^+, \tilde{H}_u^+, \tilde{W}^-, \tilde{H}_d^-)$.

The mass matrix is (in 2×2 block form):

$$\mathbf{M}_{\tilde{C}} = \begin{pmatrix} \mathbf{0} & \mathbf{X}^T \\ \mathbf{X} & \mathbf{0} \end{pmatrix}, \quad \mathbf{X} = \begin{pmatrix} M_2 & \sqrt{2} s_\beta m_W \\ \sqrt{2} c_\beta m_W & \mu \end{pmatrix} \quad (27)$$

The chargino mass matrix is diagonalized by *two* unitary 2×2 matrices \mathbf{U} and \mathbf{V} , one acting on the $+$ charged states and one acting on the $-$ charged states:

$$\begin{pmatrix} \tilde{C}_1^+ \\ \tilde{C}_2^+ \end{pmatrix} = \mathbf{V} \begin{pmatrix} \tilde{W}^+ \\ \tilde{H}_u^+ \end{pmatrix} \quad \begin{pmatrix} \tilde{C}_1^- \\ \tilde{C}_2^- \end{pmatrix} = \mathbf{U} \begin{pmatrix} \tilde{W}^- \\ \tilde{H}_d^- \end{pmatrix} \quad (28)$$

\mathbf{U} and \mathbf{V} diagonalize the mass matrix:

$$\mathbf{U}^* \mathbf{X} \mathbf{V}^{-1} = \begin{pmatrix} m_{\tilde{C}_1} & 0 \\ 0 & m_{\tilde{C}_2} \end{pmatrix} \quad (29)$$

$m_{\tilde{C}_i}^2$ are the eigenvalues of $\mathbf{X}^\dagger \mathbf{X}$.

If $m_Z \ll |\mu \pm M_2|$ then the gaugino-Higgsino mixing is small.

Then (assuming also $M_2 < |\mu|$)

$$\tilde{C}_1^\pm \approx \tilde{W}^\pm, \quad m_{C_1} \approx M_2 \quad (30)$$

$$\tilde{C}_2^+ \approx \tilde{H}_u^+, \quad \tilde{C}_2^- \approx \tilde{H}_d^-, \quad m_{C_2} \approx |\mu| \quad (31)$$

In practice the neutralino and chargino masses and mixing angles are best computed numerically.

Squarks and sleptons

In principle, all of the states with the same electric charge, R-parity, and colour can mix with each other:

$$(\tilde{u}_L, \tilde{c}_L, \tilde{t}_L, \tilde{u}_R, \tilde{c}_R, \tilde{t}_R) \rightarrow 6 \times 6 \text{ mass-squared matrix}$$

$$(\tilde{d}_L, \tilde{s}_L, \tilde{b}_L, \tilde{d}_R, \tilde{s}_R, \tilde{b}_R) \rightarrow 6 \times 6 \text{ mass-squared matrix}$$

$$(\tilde{e}_L, \tilde{\mu}_L, \tilde{\tau}_L, \tilde{e}_R, \tilde{\mu}_R, \tilde{\tau}_R) \rightarrow 6 \times 6 \text{ mass-squared matrix}$$

$$(\tilde{\nu}_e, \tilde{\nu}_\mu, \tilde{\nu}_\tau) \rightarrow 3 \times 3 \text{ mass-squared matrix}$$

Fortunately, for flavour-blind soft parameters, the inter-generation mixing is very small.

3rd family (t, b, τ) can have very different masses than first 2 families because of Yukawa and a -term contribution in RGEs.

Can also have substantial L - R mixing – see below

1st and 2nd generation sfermions end up in nearly degenerate, unmixed pairs:

$$(\tilde{e}_R, \tilde{\mu}_R), (\tilde{\nu}_e, \tilde{\nu}_\mu), (\tilde{u}_R, \tilde{c}_R), (\tilde{d}_R, \tilde{s}_R), (\tilde{u}_L, \tilde{c}_L), (\tilde{d}_L, \tilde{s}_L)$$

A second look at RGE solution for first two families (for universal m_0^2):

$$m_{Q_1}^2 = m_{Q_2}^2 = m_0^2 + K_3 + K_2 + \frac{1}{36}K_1, \quad (32)$$

$$m_{\tilde{u}_1}^2 = m_{\tilde{u}_2}^2 = m_0^2 + K_3 + \frac{4}{9}K_1, \quad (33)$$

$$m_{\tilde{d}_1}^2 = m_{\tilde{d}_2}^2 = m_0^2 + K_3 + \frac{1}{9}K_1, \quad (34)$$

$$m_{L_1}^2 = m_{L_2}^2 = m_0^2 + K_2 + \frac{1}{4}K_1, \quad (35)$$

$$m_{\tilde{e}_1}^2 = m_{\tilde{e}_2}^2 = m_0^2 + K_1. \quad (36)$$

$K_3 \gg K_2 \gg K_1$; come from SU(3), SU(2), U(1) gaugino contributions

Squarks generically heavier than sleptons.

Also a “hyperfine splitting” due D-term (sfermion)²(Higgs)² contribution:

$$\Delta_\phi = \left(T_3^\phi - Q_{EM}^\phi \sin^2 \theta_W \right) \cos 2\beta m_Z^2 \quad (37)$$

Splits members of Q , L by small amount: $\tilde{u}_L - \tilde{d}_L$, $\tilde{e}_L - \tilde{\nu}_e$

Within a single family, \tilde{q}_L - \tilde{q}_R mixing can only arise through involvement of a Higgs coupling y_q or a_q . Only important for 3rd family! E.g. for the stops:

$$\mathcal{L} = - (\tilde{t}_L^*, \tilde{t}_R^*) \mathbf{m}_t^2 \begin{pmatrix} \tilde{t}_L \\ \tilde{t}_R \end{pmatrix} \quad (38)$$

where

$$\mathbf{m}_t^2 = \begin{pmatrix} m_{Q_3}^2 + m_t^2 + \Delta_u & v(a_t \sin \beta - \mu y_t \cos \beta) \\ v(a_t \sin \beta - \mu y_t \cos \beta) & m_{\bar{u}_3}^2 + m_t^2 + \Delta_{\bar{u}} \end{pmatrix} \quad (39)$$

Diagonalize mass-squared matrix \rightarrow mass eigenstates ($m_{\tilde{t}_1}^2 < m_{\tilde{t}_2}^2$, $0 \leq \theta_{\tilde{t}} \leq \pi$):

$$\begin{pmatrix} \tilde{t}_1 \\ \tilde{t}_2 \end{pmatrix} = \begin{pmatrix} \cos \theta_{\tilde{t}} & \sin \theta_{\tilde{t}} \\ -\sin \theta_{\tilde{t}} & \cos \theta_{\tilde{t}} \end{pmatrix} \begin{pmatrix} \tilde{t}_L \\ \tilde{t}_R \end{pmatrix} \quad (40)$$

RG effects tend to make $m_{\tilde{u}_3}^2 < m_{Q_3}^2$, and both of these lighter than first 2 families. Mixing splits $m_{\tilde{t}_1}^2$ and $m_{\tilde{t}_2}^2$ even more.

Often find \tilde{t}_1 to be the lightest squark in the model!

Analogous mixing for sbottoms \tilde{b}_1, \tilde{b}_2 and staus $\tilde{\tau}_1, \tilde{\tau}_2$.

y_b, y_τ and a_b, a_τ in off-diagonal terms: mixing more important at large $\tan \beta$.

$\tilde{b}_1, \tilde{\tau}_1$ can be significantly lighter than 1st and 2nd family

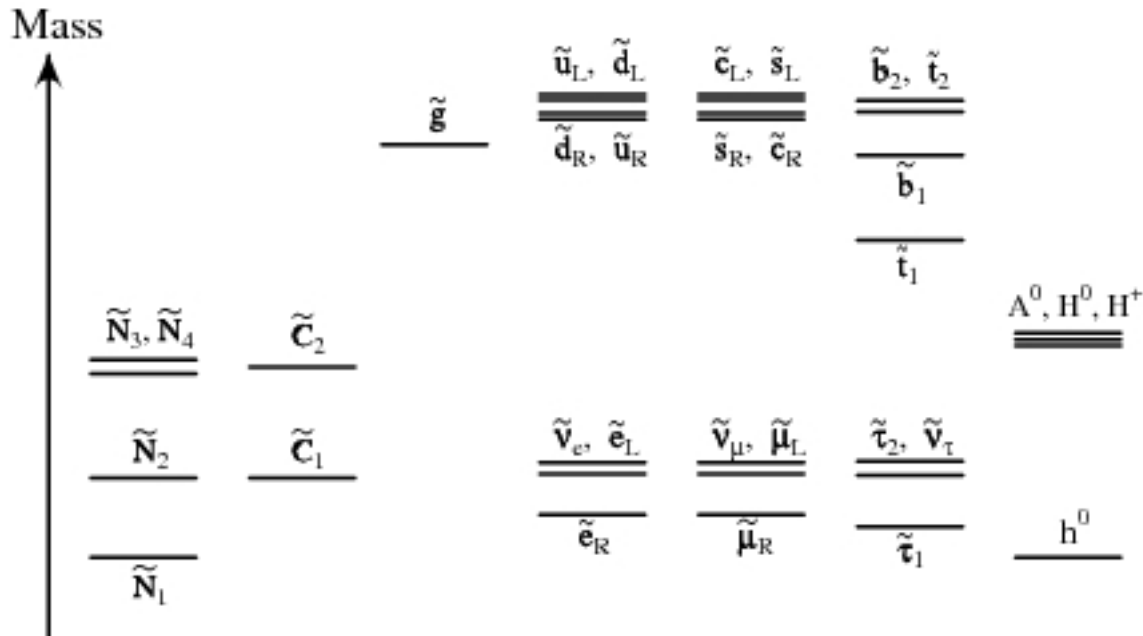
Sometimes get $\tilde{\tau}_1$ as lightest sfermion

– e.g., “stau coannihilation region” for dark matter

Summary: the particle content of the MSSM

Names	Spin	P_R	Mass Eigenstates	Gauge Eigenstates
Higgs bosons	0	+1	$h^0 \ H^0 \ A^0 \ H^\pm$	$H_u^0 \ H_d^0 \ H_u^\pm \ H_d^\mp$
squarks	0	-1	$\tilde{u}_L \ \tilde{u}_R \ \tilde{d}_L \ \tilde{d}_R$ $\tilde{s}_L \ \tilde{s}_R \ \tilde{c}_L \ \tilde{c}_R$ $\tilde{t}_1 \ \tilde{t}_2 \ \tilde{b}_1 \ \tilde{b}_2$	“ ” “ ” $\tilde{t}_L \ \tilde{t}_R \ \tilde{b}_L \ \tilde{b}_R$
sleptons	0	-1	$\tilde{e}_L \ \tilde{e}_R \ \tilde{\nu}_e$ $\tilde{\mu}_L \ \tilde{\mu}_R \ \tilde{\nu}_\mu$ $\tilde{\tau}_1 \ \tilde{\tau}_2 \ \tilde{\nu}_\tau$	“ ” “ ” $\tilde{\tau}_L \ \tilde{\tau}_R \ \tilde{\nu}_\tau$
neutralinos	1/2	-1	$\tilde{N}_1 \ \tilde{N}_2 \ \tilde{N}_3 \ \tilde{N}_4$	$\tilde{B}^0 \ \tilde{W}^0 \ \tilde{H}_u^0 \ \tilde{H}_d^0$
charginos	1/2	-1	$\tilde{C}_1^\pm \ \tilde{C}_2^\pm$	$\tilde{W}^\pm \ \tilde{H}_u^\pm \ \tilde{H}_d^\mp$
gluino	1/2	-1	\tilde{g}	“ ”
gravitino/ goldstino	3/2	-1	\tilde{G}	“ ”

A schematic sample SUSY spectrum:
 (This may or may not have anything to do with reality)



Some features of this sample spectrum:

- \tilde{N}_1 is LSP
- \tilde{t}_1 and \tilde{b}_1 are the lightest squarks
- $\tilde{\tau}_1$ is the lightest charged slepton
- Coloured particles are heavier than uncoloured particles