# 3. Heavy Quarkonia 

## 1. Spectroscopy <br> 2. em decays <br> 3. decays

# 2. The November Revolution: 

## Experimental Observation of a Heavy Particle $J \dagger$

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## and

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(Received 12 November 1974)
We report the observation of a heavy particle $J$, with mass $m=3.1 \mathrm{GeV}$ and width approximately zero. The observation was made from the reaction $p+\mathrm{Be} \rightarrow e^{+}+e^{-}+x$ by measuring the $e^{+} e^{-}$mass spectrum with a precise pair spectrometer at the Brookhaven National Laboratory's $30-\mathrm{GeV}$ alternating-gradient synchrotron

## Discovery of a Narrow Resonance in $e^{+} e^{-}$Annihilation*

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We have observed a very sharp peak in the cross section for $e^{+} e^{-} \rightarrow$ hadrons, $e^{+} e^{-}$, and possibly $\mu^{+} \mu^{*}$ at a center-of-mass energy of $3,105 \pm 0.003 \mathrm{GeV}$. The upper limit to the full width at half-maximum is 1.3 MeV .

$n_{0} *_{0}-[0 \mathrm{ev}]$


 recthe ither tan to morral rai,

pa, 1. Cross mectite woran eversy tor tol mati-






## Is Bound Charm Found?*

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We argue that the newly discovered narrow resonance at 3.1 GeV is a ${ }^{3} S_{1}$ bound state of charmed quarks and we show the consistency of this interpretation with known meson systematics. The crucial test of this notion is the existence of charmed hadrons near 2 GeV .
S. Godfrey, Carleton University

## Spectroscopy of the New Mesons*

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Center for Theoretical Physics, Massachusetts Dustitute of Technology. Cambridge, Massachusetts a21 (Received 11 December 1974)

The interpretation of the narrow boson resonances at 3.1 and 3.7 GeV as charmed quark-antiguark bound states implies the existence of other states. Some of these should be coplously produced in the radiative decays of the $3.7-\mathrm{GeV}$ resonacce. We estimate the masses and decay rates of these states and emphasize the importanee of $\gamma$-ray speetroscopy.

Two earlier papers ${ }^{1,2}$ present our case that the recently discovered ${ }^{3} 4$ and confirmed ${ }^{5}$ resonance at 3.105 GcV is the ground atate of a charmed quark bound to its antiquark, by colored gauge gluons: orthocharmonium I. More recently, a second state at 3.695 GeV has been reported ${ }^{6}$ with an estimated width of $0.5-2.7 \mathrm{MeV}$ and a partial decay rate $\sim 2 \mathrm{keV}$ into $e^{+} e^{\circ}$. We interpret this state as an $S$-wave radial excitation, orthocharmonium II, with $J^{p}=1^{-}$and $I^{G}=0^{-}$ Here are three indications of the correctness of our interpretation: (1) Much of the time, orthocharmonium II decays into orthocharmonium I and two pions. This behavior suggests that ortho charmonium II is an excited state of orthocharmonium $I^{7}{ }^{7}$ (2) The leptonic width of orthocharmonium II is about half that of orthocharmonium I, not unexpected for an excited state whose wave function at the origin is smaller. (3) Orthocharmonium II is not seen in the Brookhaven National Laboratory-Massachusetts Institute of Technology experiment, ${ }^{5}$ In a thermodynamic model, ${ }^{9}$ the production cross section of a hadron of 3.7 GeV is suppressed by $\sim 10^{-2}$ relative to that of a hadron of 3.1 GeV . Moreover, the leptonic branching ratio of orthocharmonium II is smaller than that of orthocharmonium I by a factor of 10 .
We predict the existence of other states of charmonium with masses less than 3.7 GeV , $a$


FIG. 1. Masses and radiative transitions of charmo ntum.

## Spectrum of Charmed Quark-Antiquark Bound States*

E. Eichten, K. Gottfried, T. Kinoshita, J. Kogut, K. D. Lane, and T.-M. Yan $\dagger$ Laboratory of Nuclear Studies, Cornell University, Ithaca, New York 14853 (Received 17 December 1974)

The discovery of narrow resonances at 3.1 and 3.7 GeV and their interpretation as charmed quark-antiquark bound states suggest additional narrow states between 3.0 and 4.3 GeV . A model which incorporates quark confinement is used to determine the quantum numbers and estimate masses and decay widths of these states. Their existence should be revealed by $\gamma$-ray transitions among them.

Volume 34, Number 6 PHYSical Review letters
10 Frbruary 1975


FIG. 1. The spectrum of charmonium. The vertical FIG. 1. The spectrum of charmonium. The vertical $P$ and $D$ levels are given in the text. Tbe ${ }^{2} D_{1}$ and ${ }^{1} D_{2}$ levels are not shown as their position relative to ${ }^{3} D_{1}$ ta seastive to $2^{2} s_{1}-{ }^{-} D_{1}$ mixing. Heavy lunes are allowed $E 1 \gamma$ transitions; the $2^{3} S \rightarrow 1^{1} S$ decay is a highly
suppressed $M 1$ transition. Dashed levels are unllikely to be produced or fed from above at an $e^{+} e^{-}$storage ring. Transitions among levels of an $L S$ multiplet an probably unobservable, while y transitions between states having the same value of $C=(-1)^{2+s}$ are rigor-
ously forbidden.

| Transition | $\mathrm{r}_{\boldsymbol{y}}$ | $\begin{gathered} \Gamma_{\gamma} \\ (\mathrm{keV}) \end{gathered}$ |
| :---: | :---: | :---: |
| $2^{3} S \rightarrow{ }^{3} P_{2}$ | $5_{4} \alpha k^{3}$ | 120 |
| $-^{3} P_{1}$ | ${ }^{31} H_{1} k^{3}$ | 70 |
| $\rightarrow{ }^{3} \mathrm{P}_{0}$ | $1 t_{1} \alpha k^{3}$ | 25 |
|  |  | 240 |
| ${ }^{3}{ }^{3} P_{1}=1^{3}{ }^{3} 5$ | $I_{2} \alpha h^{3}$ | 240 |
|  | ${ }_{l}^{l_{2} \alpha k^{3}}{ }_{1}{ }^{\alpha} k^{3}$ | 240 240 |
|  |  | 240 |
| $\rightarrow{ }^{{ }^{5} P_{1}}$ | $151, \alpha k^{3}$ | 110 |
|  | ${ }^{201} l_{y} \alpha k^{3}$ | ${ }_{\sim}^{150}$ |
| $2^{2} \mathrm{~S}-1^{1} \mathrm{~s}$ | $t_{\mu} k^{*}{ }^{7}$ | $\sim 1$ |

${ }^{3}$ In the second column $1 / \alpha=197, k$ is the energy of the transitton, and $I_{n}$ is a radial integral. The last oolumn is based on our wave functions and energy differences, with fine-structure splittings and $S-D$ mix-
ing ignored. ing ignored.
$P$ multiplet lies about 230 MeV below that of the $2 S$ levels. This energy difference is not very sensitive to our choice of parameters: It decreases to 160 MeV if $\alpha_{s}$ and $m_{c}$ assume the unreasonable values of 0.8 and 0.9 GeV , respectively.
(b) The c.o.g. of the lowest $D$ multiplet is 70 MeV above that of the $2 S$ levels. These $D$ levels (c)
(c) The $3 S$ level lies at $\sim 4.2 \mathrm{GeV}$. As no sharp plies that $M_{0}<4.2 \mathrm{GeV}$.

## (d) The almost GeV

## The Charmonium Spectrum

Volume 45. Number 14

Observation of an $\boldsymbol{\eta}_{\boldsymbol{c}}$ Candidate State with Mass $2978 \pm 9 \mathrm{MeV}$



Richter
Ting
Spectroscopy
convinced us that quarks were real

## "New" Spectroscopy of Mesons


S. Godfrey, Carleton University

## "New" Spectroscopy of Mesons


S. Godfrey, Carleton University

## 1. Potential Models:

- Spin independent potentials
- Relativistic corrections
- Spin dependent effects
- Coupled channel effects

Reviews:
Kwong and Rosner, Ann. Rev. Nucl. Part. Sci. 37, 325 (1987)
Buchmuller and Cooper, Adv.Ser.Direct.High Energy Phys. 1, 412 (1988)
Konigsmann, Phys. Rept. 139, 243 (1986).
Thomas as has recent review and maybe quigg?

## Mesons are composed of a quark-antiquark pair

Combine u,d,s,c,b quark and antiquark to form various mesons:
$\pi$ meson

Meson quantum numbers characterized by given JPC


Allowed:

$$
S=S_{1}+S_{2}
$$

$$
J^{P C}=0^{-+} 1^{--} 1^{+-} 0^{++} 1^{++} 2^{++} . .
$$

$$
\mathrm{J}=\mathrm{L}+\mathrm{S}
$$

$P=(-1)^{L+1}$
Not allowed: exotic combinations:

$$
J^{P C}=0^{--} 0^{+-} 1^{-+} 2^{+-} \ldots
$$

$C=(-1)^{L+S}$

### 4.1 The Spin-Independent Potential

Previously gave qualitative arguments why the spin-independent potential is linear + Coulomb

$$
V(r)=-\frac{4}{3} \frac{\alpha_{s}(r)}{r}+b r \quad b \simeq 0.18 \mathrm{GeV}^{2}
$$

We also saw how this potential is consistent with results from Lattice QCD

However, Historically this form was arrived at through trial and error (Although Appelquist and Politzer got it right in an early paper ~1975)

Emperically, the Schrodinger eqn was solved for a given potential which was modified until agreement was achieved between theory and experiment.

$$
\begin{aligned}
& M=m_{1}+m_{2}+E_{n l} \\
& {\left[\frac{p^{2}}{2 \mu}+V(r)\right] \psi=E_{n l} \psi \quad\left(\mu=\frac{m_{1} m_{2}}{m_{1}+m_{2}}\right)} \\
& {\left[\frac{\hbar^{2}}{2 \mu} \nabla^{2}+V(r)\right] \psi=E_{n l} \psi} \\
& \nabla^{2}=\frac{1}{r^{2}} \frac{\partial}{\partial r}\left(r^{2} \frac{\partial}{\partial r}\right)+\frac{1}{r^{2} \sin \theta} \frac{\partial}{\partial \theta}\left(\sin \theta \frac{\partial}{\partial \theta}\right)+\frac{1}{r^{2} \sin ^{2} \theta} \frac{\partial^{2}}{\partial \phi^{2}} \\
& \psi(r, \theta, \phi)=R(r) Y_{\ell m}(\theta, \phi) \quad U(r) \equiv r R(r) \\
& \frac{\hbar^{2}}{2 \mu} \frac{d^{2} U}{d r^{2}}+\left[V(r)+\frac{\hbar^{2}}{2 \mu} \frac{\ell(\ell+1)}{r^{2}}\right] U=E_{n \ell} U \\
& \left(U(0)=0, U^{\prime}(0)=R(0)\right) \\
& \text { S. Godfrey, Carleton University }
\end{aligned}
$$



Also $V(r)=-\frac{4}{3} \frac{\alpha_{s}}{r}+b r$ for suitable $\alpha_{s}, b$

## Lattice QCD gives q9 potential:



## Quark-antiquark Potential

For given spin and orbital angular momentum configurations \& radial excitations generate our known spectrum of light quark mesons

$$
\begin{aligned}
& H_{i j}^{\mathrm{com} f}=-\frac{4}{3} \frac{\alpha_{s}(r)}{r}+b r \\
& M=m_{1}+m_{2}+E_{n 1} \\
& {\left[\frac{p^{2}}{2 \mu}+V(r)\right] \psi=E_{n l} \psi} \\
& \text { Solve Schrodinger eqn } \\
& \text { for meson masses }
\end{aligned}
$$



Figure 21: Various $\mathbf{Q} \bar{Q}$ potentials. The potentials have been shifted to agree at $\mathrm{r}=0.5 \mathrm{fm}$. The numbers refer to the following references: 1: Martin [101], 2: Buchmüller, Grunberg and Tye [99], 3: Bhanot and Rudaz [102], 4: Cornell group [97].

From Buchmuller \& Tye PR D24, 132 (1981)

Quark potential models are strongly supported by emperical agreement with quarkonium spectroscopy and with lattice QCD
S. Godfrey, Carleton University

Could also use position of $P$-waves
Spin averaged ${ }^{3} P_{J}$ gives

$$
\bar{M}=\left(5 M_{3 P_{2}}+3 M_{3 P_{1}}+M_{3 P_{0}}\right) / 9
$$

For $c \bar{c} \bar{M}=3522 \mathrm{MeV}$

$$
\begin{aligned}
& \frac{M(2 S)-M(1 P)}{M(2 S)-M(1 S)} \begin{cases}1 / 2 & \text { H.O. }(\nu=2) \\
\simeq 1 / 4 & \text { for } \nu=0 \\
0 & \text { Coulomb }(\nu=-1)\end{cases} \\
& c \bar{c} \Rightarrow \nu \simeq 0.15
\end{aligned}
$$

## Spin-dependent potentials:

Generally expect spin-dependent Interactions:

$$
\vec{S}_{1} \cdot \vec{S}_{2} \quad \vec{L} \cdot \vec{S} \quad S_{12}
$$

Start by looking at spin-dependent interactions of QED in hydrogen atom

Spin-Orbit: electron sees the proton circling around
-The orbital motion creates a magnetic field at the centre:

- In terms of $L=m v r$

$$
B=\frac{e v}{c r^{2}}
$$

$$
\vec{B}=\frac{e}{m c r^{3}} \vec{L}
$$

-The spinning electron constitutes a magnetic dipole

$$
\vec{\mu}=-\frac{e}{m c} \vec{S}
$$

-The interaction energy is

$$
W=-\vec{\mu} \cdot \vec{B}
$$

More rigorously (derived as a succession of infinitesimal Lorentz transformations) leads to the Thomas precession with a factor of $1 / 2$

$$
\begin{gathered}
\Delta H_{S . O}=\frac{e^{2}}{2 m^{2} c^{2} r^{3}} \vec{L} \cdot \vec{S} \\
\overrightarrow{J^{2}}=\vec{L}^{2}+\vec{S}^{2}+2 \vec{L} \cdot \vec{S} \\
\Rightarrow \vec{L} \cdot \vec{S}=\frac{1}{2}\left[\vec{J}^{2}-\vec{L}^{2}-\vec{S}^{2}\right] \\
=\frac{1}{2}[J(J+1)-L(L+1)-S(S+1)] \\
\text { For } \begin{cases}{ }^{3} P_{2} & \vec{L} \cdot \vec{S}=1 \\
{ }_{3} P_{1} & \vec{L} \cdot \vec{S}=-1 \\
{ }^{3} P_{0} & \vec{L} \cdot \vec{S}=-2\end{cases}
\end{gathered}
$$

Hyperfine: Again in hydrogen, the proton has dipole moment:

$$
\vec{\mu}_{P}=\gamma_{P} \frac{e}{m_{P c}} \vec{S}_{P}\left(\gamma_{P}=2.73\right)
$$

The magnetic dipole has a field:

$$
\vec{B}(\vec{r})=\underbrace{\frac{1}{r^{3}}\left[\frac{3(\vec{\mu} \cdot \vec{r}) \vec{r}}{r^{2}}-\vec{\mu}\right]}_{r>a}+\underbrace{\frac{8 \pi}{3} \vec{\mu}}_{r<a}
$$



The energy of the electon in the presence of $\mu_{\mathrm{i}}$

$$
\left.\Delta H_{S S}=\frac{\gamma_{P} e^{2}}{m m_{P} c^{2}}\left\{\frac{1}{r^{3}}\left[3\left(\vec{S}_{P} \cdot \hat{r}\right)\left(\vec{S}_{e} \cdot \hat{r}\right)-\vec{S}_{P} \cdot \vec{S}_{e}\right)\right]+\frac{8 \pi}{3} \vec{S}_{P} \cdot \vec{S}_{e} \delta^{3}(\vec{r})\right\}
$$

Gives rise to the hyperfine structure of hydrogen

$$
s \quad+1 / 4{ }^{3} s_{1} \quad \vec{S}_{1} \cdot \vec{S}_{2}=\frac{1}{2}\left[\vec{S}^{2}-\vec{S}_{1}^{2}-\vec{S}_{2}^{2}\right]=\frac{1}{2}\left[s(s+1)-\frac{3}{2}\right]
$$

$$
-3 / 4 \text { ts. } 21 \mathrm{~cm} \text { line in hydrogen }
$$

## One can take this over to 1-gluon interaction of QCD:

$$
\begin{gathered}
\left.\Delta H_{i j}^{h y p}=-\frac{\alpha_{s}(r)}{m_{i} m_{j}}\left\{\frac{8 \pi}{3} \vec{S}_{i} \cdot \vec{S}_{j} \delta^{3}\left(\vec{r}_{i j}\right)+\frac{1}{r_{i j}^{3}}\left[3\left(\vec{S}_{i} \cdot \hat{r}_{i j}\right)\left(\vec{S}_{j} \cdot \hat{r}_{i j}\right)-\vec{S}_{i} \cdot \vec{S}_{j}\right)\right]\right\} \vec{F}_{i} \cdot \vec{F}_{j} \\
\Delta H_{i j}^{\text {S.O.(c.m.) }}=-\frac{\alpha_{s}(r)}{r_{i j}^{3}}\left(\frac{1}{m_{i}}+\frac{1}{m_{j}}\right)\left(\frac{\vec{S}_{i}}{m_{i}}+\frac{\vec{S}_{j}}{m_{j}}\right) \cdot \vec{L} \vec{F}_{i} \cdot \vec{F}_{j} \\
\Delta H_{i j}^{S . O .(T P)}=-\frac{1}{2 r_{i j}} \frac{\partial V(r)}{\partial r_{i j}}\left(\frac{\vec{S}_{i}}{m_{i}^{2}}+\frac{\vec{S}_{j}}{m_{j}^{2}}\right) \cdot \vec{L} \vec{F}_{i} \cdot \vec{F}_{j} \\
\text { For mesons }\left\langle\vec{F}_{i} \cdot \vec{F}_{j}\right\rangle=-\frac{4}{3}
\end{gathered}
$$

Systematic treatment starts with Wilson loop
Eichten and Feinberg, PR D23, 2724 (1981) Gromes, Yukon Advanced Study Inst.

- Expanding in $1 / m_{Q}$ write spin-dependent Hamiltonian in terms of static potential and correlation functions of colour electric and magnetic fields
-With some assumptions one obtains:

$$
\begin{aligned}
V_{\text {spin }}(r)= & \frac{1}{m^{2}}\left(\frac{-k}{2 r}+\frac{2 \alpha_{s}}{3 r^{3}}\right) \vec{L} \bullet \vec{S} \\
& +\frac{1}{m^{2}} \frac{4 \alpha_{s}}{3 r^{3}} S_{12}+\frac{1}{m^{2}} \frac{32 \pi \alpha_{s}}{9} \delta^{3}(\vec{r}) \vec{S}_{1} \bullet \vec{S}_{2}
\end{aligned}
$$

Which corresponds to short range vector and long range scalar exchange

## Observation of ${ }^{1} P_{1}$ states is important test

## Spin-dependent potentials:

- Need some sort of reduction to find spin dependent terms
- Depends on Lorentz nature of potential
we find phenomenologically
short range Lorentz Vector 1-gluon exchange
+ long range Lorentz scalar confining potential
- Use Breit-Fermi Hamiltonian
- Spin-dependent interactions are (v/c) ${ }^{2}$ corrections

Spin-spin interactions:

$$
\begin{aligned}
& H_{i j}^{h h_{j p}}=\frac{4 \alpha_{\alpha}(r)}{3 m_{i} m_{j}}\left\{\frac{8 \pi}{3} \vec{S}_{i} \cdot \vec{S}_{j} \delta^{3}\left(\vec{r}_{i j}\right)+\frac{1}{r_{i j}^{3}}\left[\frac{3 \vec{S}_{i} \cdot \vec{r}_{i j} \vec{S}_{j} \cdot \vec{r}_{i j}}{r_{i j}^{2}}-\vec{S}_{i} \cdot \vec{S}_{j}\right]\right\} \\
& \vec{S}_{1} \cdot \vec{S}_{2}=\frac{1}{2}\left[S^{2}-S_{1}^{2}-S_{2}^{2}\right]=\frac{1}{2}\left[s(s+1)-\frac{3}{2}\right]
\end{aligned}
$$



Spin-orbit interactions:

$$
\begin{aligned}
& H_{i j}^{\text {s.o. }(\mathrm{cm})}=\frac{4 \alpha_{s}(r)}{3 r_{i j}^{3}}\left(\frac{1}{m_{i}}+\frac{1}{m_{j}}\right)\left(\frac{\vec{S}_{i}}{m_{i}}+\frac{\vec{S}_{j}}{m_{j}}\right) \cdot \vec{L} \\
& H_{i j}^{\text {s.o. }(t p)}=\frac{-1}{2 r_{i j}} \frac{\partial V(r)}{\partial r_{i j}}\left(\frac{\vec{S}_{i}}{m_{i}^{2}}+\frac{\vec{S}_{j}}{m_{j}^{2}}\right) \cdot \vec{L}
\end{aligned}
$$

$$
\begin{aligned}
& { }^{3} P_{2}: \vec{L} \cdot \vec{S}=1 \\
& { }^{3} P_{1}: \vec{L} \cdot \vec{S}=-1 \\
& { }^{3} P_{0}: \vec{L} \cdot \vec{S}=-2
\end{aligned}
$$



Let us examine the spin-dependent splittings in charmonium

- Using H.O. wavefunctions simplifies the calculations
- Fitting the oscillator parameter to the r.m.s. radii of exact solutions is a good approximation:

$$
\begin{array}{ll}
\psi_{1 S}=\frac{2}{\pi^{1 / 4}} \beta^{3 / 2} e^{-\beta^{2} r^{2} / 2} Y_{00} & \left\langle r^{2}\right\rangle_{1 S}=\frac{3}{2} \frac{1}{\beta^{2}}=2.5 \Rightarrow \beta=0.77 \\
\psi_{2 S}=\sqrt{\frac{8}{3}} \frac{\beta^{3 / 2}}{\pi^{1 / 4}}\left(\frac{3}{2}-\beta^{2} r^{2}\right) e^{-\beta^{2} r^{2} / 2} Y_{00} & \left\langle r^{2}\right\rangle_{2 S}=\frac{7}{2} \frac{1}{\beta^{2}}=11 \Rightarrow \beta=0.564 \\
\psi_{1 P}=\sqrt{\frac{8}{3}} \frac{\beta^{5 / 2} r}{\pi^{1 / 4}} e^{-\beta^{2} r^{2} / 2} Y_{1 m} & \left\langle r^{2}\right\rangle_{1 P}=\frac{5}{2} \frac{1}{\beta^{2}} \simeq 7 \Rightarrow \beta=0.598 \\
& \langle 1 / r\rangle_{1 P}=\frac{4}{3} \frac{\beta}{\pi^{1 / 2}}=0.45 \\
& \left\langle 1 / r^{3}\right\rangle_{1 P}=\frac{4}{3} \frac{\beta^{3}}{\pi^{1 / 2}}=0.16
\end{array}
$$

## Hyperfine Effects:

$$
\begin{aligned}
H_{i j}^{h y p} & =\frac{32 \pi}{9} \frac{\alpha_{s}}{m^{2}} \vec{S}_{1} \cdot \vec{S}_{2} \delta^{3}\left(r_{i j}\right) \\
\vec{S}_{1} \cdot \vec{S}_{2} & =\frac{1}{2}[s(s+1)-3 / 2] \\
& \Rightarrow \begin{cases}\left.\left.\left\langle{ }^{3} S_{1}\right| \vec{S}_{1} \cdot \vec{S}_{2}\right|^{3} S_{1}\right\rangle=+1 / 4 \\
\left\langle{ }^{1} S_{0}\right| \vec{S}_{1} \cdot \vec{S}_{2}\left|{ }^{1} S_{0}\right\rangle=-3 / 4\end{cases}
\end{aligned}
$$

$$
\begin{aligned}
\therefore M\left({ }^{3} S_{1}\right)-M\left({ }^{1} S_{0}\right) & =\frac{32 \pi}{9} \frac{\alpha_{s}}{m^{2}}\left\langle\delta^{3}\left(r_{i j}\right)\right\rangle \\
& =\frac{32 \pi}{9} \frac{\alpha_{s}}{m^{2}}|\psi(0)|^{2} \\
& =\frac{32 \pi}{9} \frac{\alpha_{s}}{m^{2}} \frac{\beta^{3}}{\pi^{3 / 2}}
\end{aligned}
$$

$$
\begin{equation*}
=115 \mathrm{MeV}\left(\text { where } \beta=0.77 \mathrm{GeV}, \alpha_{s}=0.32, m_{c}=1.6 \mathrm{GeV}\right) \tag{1}
\end{equation*}
$$

vs 115 MeV from experiment

$$
M\left(2^{3} S_{1}\right)-M\left(2^{1} S_{0}\right)=67 \mathrm{MeV}
$$

Fine Structure:
We can write the ${ }^{3 P_{\mathrm{J}}}$ Masses as:

$$
\begin{aligned}
M= & M(1 P)+a\langle\vec{L} \cdot \vec{S}\rangle+b\left\langle S_{12}\right\rangle \\
M\left({ }^{3} P_{2}\right)= & M(1 P)+a-\frac{2}{5} b=3556 \\
M\left({ }^{3} P_{1}\right)= & M(1 P)-a-2 b=3511 \\
M\left({ }^{3} P_{0}\right)= & M(1 P)-2-4 b=3415 \\
& M(1 P)=3525
\end{aligned}
$$

Lorentz Vector 1-gluon exchange gives:

$$
\begin{aligned}
a & =\frac{3}{2 m^{2}} \frac{4}{3} \frac{\alpha_{s}}{r^{3}}=40 \mathrm{MeV} \\
b & =\frac{1}{4 m^{2}} \frac{4}{3} \frac{\alpha_{s}}{r^{3}}=7 \mathrm{MeV}
\end{aligned}
$$

## ${ }^{1} P_{1}$ vs ${ }^{3} P_{\text {cog }}$ mass - distinguish models

## - Important to distinguish models

- In QM triplet-singlet splittings test
- the Lorentz nature of the confining potential
- Relativistic effects
-important validation of
- lattice QCD calculations
- NRQCD calculations


Observation of ${ }^{1} P_{1}$ states is an important test of theory

## Wide variation of theoretical predictions:



Quark Potential Models with 1-gluon exchange:

$$
H_{q \bar{q}}^{h y p}=\frac{32 \pi}{9} \frac{\alpha_{s}}{m_{q} m_{\bar{q}}} \vec{S}_{q} \cdot \vec{S}_{\bar{q}} \delta^{3}(\vec{r})
$$

$\delta$ function is short range but smeared by relativistic effects modeled by a Gaussian.
-gives $M\left({ }^{3} P_{c o g}\right)>M\left({ }^{1} P_{1}\right)$
Godfrey \& Isgur, PR D32, 189 (1985)
wide variation in predictions indicates need for experimental data

## Decays and Transitions



- To calculate Decays and Transitions we need to calculate hadronic matrix elements.
- Define a "Mock" meson which we equate with the wavefucntion of the physical meson
$\left.\left|M(\vec{K})=\sqrt{2 E_{M}} \int d^{3} p \Phi(\vec{p}) \chi_{s \bar{s}} \phi_{q \bar{q}} \phi_{\text {colour }}\right| q\left(\frac{m_{q}}{m_{q}+m_{\bar{q}}} \vec{K}+p, s\right) \bar{q}\left(\frac{m_{\bar{q}}}{m_{q}+m_{\bar{q}}} \vec{K}-p, \bar{s}\right)\right\rangle$

There are two generic types of matrix elements:

$$
\begin{gathered}
\langle 0| A\left|M_{i}\right\rangle \text { like in } J / \psi \rightarrow e^{+} e^{-} \\
\left\langle M_{f}\right| A\left|M_{i}\right\rangle \text { like in } \chi_{c 2} \rightarrow J / \psi+\gamma
\end{gathered}
$$

A is some sort of transition operator like:

$$
j_{e m}^{\mu}=\bar{q} \gamma^{\mu} q
$$

## Crystal Ball



Fig. 5. Inclusive photon spectrum at the $\psi^{\prime}$ obtained by the Crystal Ball experiment. Note that the logarithmic energy scale yields bin sizes approximately proportional to photon energy resolution. The numbers over the spectrum correspond to the expected radiative transitions shown in the enectrum incet
S. Godfrey, Carleton University


Fig. 4. - The $\mathrm{b} \overline{\mathrm{b}}$ level diagram showing transitions between states. We use the familiar spectroscopic notation $n^{2 S+1} L_{J}$, where $n$ is the principal quantum number (with the convention that $n$ is one plus the number of nodes in the wavefunction), and $L, S$, and $J$ are the orbital angular momentum, total spin, and total angular momentum. The parity and $C$-parity are given by $P=(-)^{L+1}$ and $C=(-)^{L+S}$. Note that not all states and transitions shown have been


Figure 11: ARGUS [74] $\Upsilon(2 S) \rightarrow \gamma+$ hadrons with $\gamma \rightarrow \mathrm{e}^{+} \mathrm{e}^{-}$.


Figure 12: C'LEO [75] $\Upsilon(2 S)-\gamma+$ hadrons with $\gamma \rightarrow \mathrm{e}^{+} \mathrm{e}^{-}$. The data do not re S. Godfrey, Carleton Univlines and it is not included in the fit shown.

## Photon Transitions in $Y(2 S)$ and $Y(3 S)$ Decays



FIG. 2. Fit to the $\Upsilon(2 S) \rightarrow \gamma \chi_{b J}(1 P)(J=2,1,0)$ photon lines in the data. The points represent the data (top plot). Statistical errors on the data are smaller than the point size. The solid line represents the fit. The dashed line represents total fitted background. The background subtracted data (points with error bars) are shown at the bottom. The solid line represents the fitted photon lines together. The dashed lines show individual photon lines.


FIG. 3. Fit to the $\Upsilon(3 S) \rightarrow \gamma \chi_{b J}(2 P)(J=2,1,0)$ photon lines in the data. See caption of Fig. 2 for the description. Small solid line peaks in the bottom plot show the $\chi_{b J}(2 P) \rightarrow \gamma \Upsilon(1 D)$ and $\mathrm{Y}(2 S) \rightarrow \gamma \chi_{h I}(1 P)$ contributions.

## Radiative (e.m.) Transitions

Same physics as in atomic and nuclear systems An e.m. transition is described by:


For 2 body decay $M_{i} \rightarrow M_{f} \gamma$

$$
\begin{aligned}
d \Gamma= & \frac{(2 \pi)^{4} \delta^{4}\left(P_{f}+p_{\gamma}-p_{i}\right)}{2 M_{i}}\left|M_{f i}\right|^{2} \frac{d^{3} p_{f}}{(2 \pi)^{3}\left(2 E_{f}\right)} \frac{d^{3} p_{\gamma}}{(2 \pi)^{3}\left(2 E_{\gamma}\right)} \\
\Gamma= & \frac{1}{2 \pi M^{2}}\left|M_{i f}\right|^{2} p \\
& \text { where } p=\frac{\left(M_{i}^{2}-M_{f}^{2}\right)}{2 M_{i}} \\
= & \frac{\left|M_{i f}\right|^{2}}{8 \pi M}\left(1-M_{f}^{2} / M_{i}^{2}\right) \\
\frac{d \Gamma}{d \cos \theta}= & \frac{\left|M_{i f}\right|^{2}}{16 \pi^{2} M_{i}}\left(1-M_{f}^{2} / M_{i}^{2}\right)=\frac{\left|M_{i f}\right|^{2}}{8 \pi^{2} M_{i}^{2}} k_{\gamma}
\end{aligned}
$$

## Start with E1 Transitions:

$\frac{p^{2}}{2 m} \rightarrow \frac{(\vec{p}-e \vec{A})^{2}}{2 m}=\frac{p^{2}}{2 m}-\frac{e \vec{p} \cdot \vec{A}}{2 m}-\frac{e \vec{A} \cdot \vec{p}}{2 m}+e^{2} \frac{\vec{A}^{2}}{2 m}$
$p^{2} / 2 m$ is the original kinetic energy term
drop higher order $e^{2} \vec{A}^{2}$ terms
Interested in:

$$
\begin{aligned}
& H_{I}=-\frac{e}{2 m}(\vec{A} \cdot \vec{p}+\vec{p} \cdot \vec{A}) \\
& \vec{A}(x)=\frac{1}{\sqrt{2 \omega}} \vec{\epsilon}(\vec{k}) e^{i \vec{k} \cdot \vec{x}} \\
& e^{i \vec{k} \cdot \vec{x}} \simeq 1+i \vec{k} \cdot \vec{x}+\ldots \\
& \text { in the long wavelength limit } \frac{1}{k} \gg r \\
& \Rightarrow \vec{A}(x) \simeq \frac{1}{\sqrt{2 \omega}} \vec{\epsilon}(\vec{k}) \\
& H_{I}=-\frac{e}{2 m}(\vec{\epsilon} \cdot \vec{p}+\vec{p} \cdot \vec{\epsilon})
\end{aligned}
$$

To evaluate $\langle A| \vec{p}|B\rangle \cdot \vec{\epsilon}$
Start with $\left[p_{i}, r_{j}\right]=-i \delta_{i j}$

$$
\begin{aligned}
\Rightarrow & {\left[\vec{p}^{2}, r_{j}\right]=p_{i}\left[p_{i}, r_{j}\right]+\left[p_{i}, r_{j}\right] p_{i}=-2 i p_{j} } \\
\langle A| p_{i}|B\rangle= & i\langle A|\left[\vec{p}^{2} / 2, r_{j}\right]|B\rangle \\
= & i \mu\langle A|\left[H, r_{j}\right]|B\rangle \\
& \left(H=p^{2} / 2 \mu+V(r) \text { but }[V(r), r]=0\right) \\
= & i \mu\langle A| H r_{j}-r_{j} H|B\rangle \\
= & i \mu\left(E_{A}-E_{B}\right)\langle A| r_{j}|B\rangle \\
= & i \frac{m}{2} \omega\langle A| r_{j}|B\rangle \\
\langle A| H_{I}|B\rangle= & -\frac{i e m \omega}{2 m}\langle A| r_{i}|B\rangle \epsilon_{i} \\
= & -\frac{i e \omega}{2}\langle A| r_{i}|B\rangle \epsilon_{i} \\
= & -\frac{i e \omega}{2}\langle A| \vec{r}|B\rangle \cdot \vec{\epsilon}
\end{aligned}
$$

There are two methods for evaluating the matrix element Method 1:

The sum over final polarizations is:

$$
\sum_{p o l} \vec{\epsilon}_{i}(k) \vec{\epsilon}^{*}(k)=\delta_{i j}-k_{i} k_{j} / \vec{k}^{2}
$$

So:

$$
\left.\left.\left.\sum_{p o l}\left|\langle B| H_{I}\right| A\right\rangle\left.\right|^{2}=\left.\omega^{2} e^{2} Q^{2}\{|\langle B| \vec{r}| A\rangle\right|^{2}-|\langle B| \vec{r} \cdot \hat{k}| A\right\rangle\left.\right|^{2}\right\}
$$

Averaging over directions:

$$
\left.=\omega^{2} e^{2} Q^{2} \frac{2}{3}|\langle B| \vec{r}| A\right\rangle\left.\right|^{2}
$$

Start with ${ }^{3} \mathrm{P}_{\mathrm{J}} \rightarrow{ }^{3} \mathrm{~S}_{1}$

- The orbital angular momentum is zero in the final state
-We may choose any $J_{z}$ since we averaged over the photon directions

Convenient to choose $\mathrm{J}_{\mathrm{Z}}=\mathrm{J}$

Start by writing down the meson wavefunction: $|M\rangle=\sqrt{2 M} \psi(r)$ where $\sqrt{2 M}$ is introduced to normalize the wavefunction when integrating over relativistic phase space.

$$
{ }^{3} P_{2}\left(J_{z}=2\right):\left|J=J_{Z}=2\right\rangle=\left|L=L_{Z}=1\right\rangle \otimes\left|S=S_{Z}=1\right\rangle=\left|Y_{11} \uparrow \uparrow\right\rangle
$$

Only $J_{Z}^{\prime}=S_{Z}^{\prime}=1$ contributes since $H_{I}$ does not flip spin.

$$
\begin{aligned}
& \langle f| \vec{r}|i\rangle=\langle f| r|i\rangle \int\left\langle Y_{00} \uparrow \uparrow\right| \sqrt{\frac{4 \pi}{3}} Y_{1-1}\left|Y_{11} \uparrow \uparrow\right\rangle d \Omega=\langle f| r|i\rangle \sqrt{\frac{1}{3}} \\
& \text { where }\langle f| r|i\rangle=\int r^{2} d r R_{f}(r) r R_{i}(r) \sqrt{2 M_{i}} \sqrt{2 M_{f}} \\
& \begin{aligned}
& \Gamma\left({ }^{3} P_{2}\right.\left.\rightarrow{ }^{3} S_{1}\right)=\frac{1}{8 \pi M^{2}}\left|M_{i f}\right|^{2} \omega \\
&\left.\quad=\frac{\omega}{8 \pi M^{2}} \omega^{2} e^{2} Q^{2}|\langle f| r| i\right\rangle\left.\right|^{2}\left(s M_{i}\right)\left(2 M_{f}\right) \times \frac{2}{3} \times \frac{1}{3} \\
&\left.\quad=\frac{4 \pi \alpha \omega^{3} e_{q}^{2}}{8 \pi} \frac{8}{9}|\langle f| r| i\right\rangle\left.\right|^{2}\left(\frac{M_{i} M_{f}}{M_{i} M_{i}}\right) \\
&\left.\quad=\frac{4}{9} \alpha \omega^{3} e_{q}^{2}|\langle f| r| i\right\rangle\left.\right|^{2}\left(\frac{M_{f}}{M_{i}}\right)
\end{aligned}
\end{aligned}
$$

For ${ }^{3} P_{1} \rightarrow{ }^{3} S_{1}$
$\left|J=J_{Z}=1\right\rangle=\frac{1}{\sqrt{2}}\left|Y_{11} \frac{1}{\sqrt{2}}(\uparrow \downarrow+\downarrow \uparrow)-Y_{10} \uparrow \uparrow\right\rangle$
so that

$$
\begin{aligned}
& \left\langle Y_{00}\right| \vec{r}\left|J=J_{Z}=1\right\rangle=\frac{1}{\sqrt{2}}\left\langle Y_{00}\right| \vec{r}\left|Y_{11}\right\rangle-\frac{1}{\sqrt{2}}\left\langle Y_{00}\right| \vec{r}\left|Y_{10}\right\rangle \\
& \quad=\left[\frac{1}{\sqrt{2}}\left\langle Y_{00}\right| \frac{1}{\sqrt{3}}\left(\frac{-\hat{x}+i \hat{y}}{\sqrt{2}}\right)\left|Y_{11}\right\rangle-\frac{1}{\sqrt{2}}\left\langle Y_{00}\right| \frac{1}{\sqrt{3}} \hat{z}\left|Y_{10}\right\rangle\right]\langle 1 S| r|1 P\rangle \\
& \left.\left.\Rightarrow\left|\left\langle^{3} S_{1}\right| \vec{r}\right|^{3} P_{1}\right\rangle\left.\right|^{2}=\left[\frac{1}{2} \frac{1}{3}+\frac{1}{2} \frac{1}{3}\right]|\langle 1 S| r| 1 P\right\rangle\left.\right|^{2} \\
& { }^{3} P_{0} \rightarrow{ }^{3} S_{1}
\end{aligned}
$$

$$
\left|J=J_{Z}=0\right\rangle=\sqrt{\frac{1}{3}}\left|Y_{11} \downarrow \downarrow-Y_{10} \sqrt{\frac{1}{2}}(\uparrow \downarrow+\downarrow \uparrow)+Y_{1-1} \uparrow \uparrow\right\rangle
$$

$$
\text { resulting in } \left.\left.\left|\left\langle^{3} S_{1}\right| \vec{r}\right|^{3} P_{0}\right\rangle\left.\right|^{2}=\left[\frac{1}{3} \frac{1}{3}+\frac{1}{3} \frac{1}{3}+\frac{1}{3} \frac{1}{3}\right]|\langle 1 S| r| 1 P\right\rangle\left.\right|^{2}
$$

Summarizing all these results we obtain:

$$
\begin{aligned}
& \left.\Gamma\left({ }^{3} P_{2} \rightarrow{ }^{3} S_{1} \gamma\right)=\frac{\omega^{3} e^{2} Q^{2}}{3 \pi} \frac{1}{3}|\langle 1 S| r| 1 P\right\rangle\left.\right|^{2} \\
& \left.\Gamma\left({ }^{3} P_{1} \rightarrow{ }^{3} S_{1} \gamma\right)=\frac{\omega^{3} e^{2} Q^{2}}{3 \pi}\left\{\frac{1}{2} \frac{1}{3}+\frac{1}{2} \frac{1}{3}\right\} \frac{1}{3}|\langle 1 S| r| 1 P\right\rangle\left.\right|^{2} \\
& \left.\Gamma\left({ }^{3} P_{0} \rightarrow{ }^{3} S_{1} \gamma\right)=\frac{\omega^{3} e^{2} Q^{2}}{3 \pi}\left\{\frac{1}{3} \frac{1}{3}+\frac{1}{3} \frac{1}{3}+\frac{1}{3} \frac{1}{3}\right\}|\langle 1 S| r| 1 P\right\rangle\left.\right|^{2}
\end{aligned}
$$

Comparing these expressions we see that in all cases

$$
\left.\Gamma\left({ }^{3} P_{J} \rightarrow{ }^{3} S_{1} \gamma\right)=\frac{4 \alpha \omega^{3} Q^{2}}{9}|\langle 1 S| r| 1 P\right\rangle\left.\right|^{2}
$$

Similarly we obtain:

$$
\left.\Gamma\left({ }^{3} S_{1} \rightarrow{ }^{3} P_{J} \gamma\right)=\frac{4 \alpha \omega^{3} Q^{2}(2 J+1)}{27}|\langle 1 S| r| 1 P\right\rangle\left.\right|^{2}
$$

## Let us return to our effective wavefunctions:

$$
\begin{array}{rlr}
\psi_{1 S} & =\frac{2}{\pi^{1 / 4}} \beta^{3 / 2} e^{-\beta^{2} r^{2} / 2} Y_{00} & \beta=0.77 \mathrm{GeV} \\
\psi_{1 P} & =\sqrt{\frac{8}{3}} \frac{\beta^{5 / 2} r}{\pi^{1 / 4}} e^{-\beta^{2} r^{2} / 2} Y_{1 m} & \beta=0.598 \mathrm{GeV}
\end{array}
$$

This gives:

$$
\begin{aligned}
& \left\langle\psi_{1 S}\right| r\left|\psi_{1 P}\right\rangle=\frac{2}{\pi^{1 / 4}} \sqrt{\frac{8}{3}} \frac{1}{\pi^{1 / 4}} \beta_{S}^{3 / 2} \beta_{P}^{5 / 2} \int r^{4} e^{-\left(\beta_{S}^{2}+\beta_{P}^{2}\right) r^{2} / 2} d r \\
& =\sqrt{\frac{8}{3}} 15 \frac{\beta_{S}^{3 / 2} \beta_{P}^{5 / 2}}{\left(\beta_{S}^{2}+\beta_{P}^{2}\right)^{5 / 2}} \\
& =5.2 \mathrm{GeV}^{-1} \\
& \Rightarrow \Gamma\left({ }^{3} P_{2} \rightarrow{ }^{3} S_{1} \gamma\right)=0.59 \mathrm{MeV} \quad \text { vs } \quad \Gamma^{e x p t}=0.351_{-.14}^{+.2} \mathrm{MeV} \\
& \Gamma\left({ }^{3} P_{1} \rightarrow{ }^{3} S_{1} \gamma\right)=\mathrm{MeV} \quad \text { vs } \quad \Gamma^{e x p t}<0.355 \mathrm{MeV}
\end{aligned}
$$

## 3. E1 transitions

E1 decays sensitive to nodes in wavefunction
radiative transitions tests internal structure


## Including relativistic corrections corresponds to using

 eigenfunctions and eigenvalues of the Breit-Fermi Hamiltonian (Siegert's theorem)|  | $<2 P\|r\| 3 S>\mid$ |  | $\langle 1 P\| r\|2 S\rangle$ |  | $<1 P\|r\| 3 S>$ |  | $\begin{aligned} & \hline\langle 1 S\| r\|2 P\rangle \\ & \|\langle 2 S\| r\| 2 P\rangle \mid \\ & \hline \end{aligned}$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | $\mathrm{GeV}^{-1}$ |  | $\mathrm{GeV}^{-1}$ |  | $\mathrm{GeV}^{-1}$ |  |  |  |
| DATA | $2.7 \pm 0.2$ |  | $1.9 \pm 0.2$ |  | $0.050 \pm 0.006$ |  | $0.096 \pm 0.005$ |  |
|  | World Average |  |  |  | This measurement |  |  |  |
| Model | NR | rel | NR | rel | NR | rel | NR | rel |
| Kwong, Rosner [13] | 2.7 |  | 1.6 |  | 0.023 |  | 0.13 |  |
| Fulcher [14] | 2.6 |  | 1.6 |  | 0.023 |  | 0.13 |  |
| Büchmuller et al.[15] | 2.7 |  | 1.6 |  | 0.010 |  | 0.12 |  |
| Moxhay,Rosner [16] | 2.7 | 2.7 | 1.6 | 1.6 | 0.024 | 0.044 | 0.13 | 0.15 |
| Gupta et al.[17] | 2.6 |  | 1.6 |  | 0.040 |  | 0.11 |  |
| Gupta et al.[18] | 2.6 |  | 1.6 |  | 0.010 |  | 0.12 |  |
| Fulcher [19] | 2.6 |  | 1.6 |  | 0.018 |  | 0.11 |  |
| Danghighian et al.[20] | 2.8 | 2.5 | 1.7 | 1.3 | 0.024 | 0.037 | 0.13 | 0.10 |
| McClary,Byers [21] | 2.6 | 2.5 | 1.7 | 1.6 |  |  | 0.15 | 0.13 |
| Eichten et al.[22] | 2.6 |  | 1.7 |  | 0.110 |  | 0.15 |  |
| Grotch et al.[23] | 2.7 | 2.5 | 1.7 | 1.5 | 0.011 | 0.061 | 0.13 | 0.19 |

Tomasz Skwarnicki, Syracuse U. ICHEP, Amsterdam July,2002

Relativistic effects gives differences between E1 matrix elements:

$$
\begin{aligned}
& \langle 2 P| r|3 S\rangle=2.7 \pm 0.2 \mathrm{GeV}^{-1} \\
& \left\langle 2^{3} P_{2} \mid r 3^{3} S_{1}\right\rangle \approx-2.4 \mathrm{GeV}^{-1} \\
& \left\langle 2^{3} P_{1} \mid 3^{3} S_{1}\right\rangle \approx-2.3 \mathrm{GeV}^{-1} \\
& \left\langle 2^{3} P_{0}\right| r\left|3^{3} S_{1}\right\rangle \approx-2.2 \mathrm{GeV}^{-1} \\
& \langle 1 P| r|2 S\rangle \pm 1.9 \pm 0.2 \mathrm{GeV}^{-1} \\
& \left.\left.\left\langle 1^{3}\right||r|\right|^{3} S_{1}\right\rangle \approx-1.5 \mathrm{GeV}^{-1} \\
& \left\langle 1^{3} P_{1}\right| r\left|2^{3} S_{1}\right\rangle \approx-1.4 \mathrm{GeV}^{-1} \\
& \left\langle 1^{3} P_{0}\right| r\left|2^{3} S_{1}\right\rangle \approx-1.3 \mathrm{GeV}^{-1}
\end{aligned}
$$



see also McClary and Byers, PR D28, 1692 (1983)
S. Godfrey, Carleton University

| $\|$$\mid\langle 1 P\| r\|3 S>\|$  <br> $\mathrm{GeV}^{-1}$  <br> This mea!  <br> NR rel <br> $0.050 \pm 0.006$  <br> 0.023  <br> 0.010  <br> 0.024 0.044 <br> 0.040  <br> 0.010  <br> 0.018  <br> 0.024 0.037 <br> 0.110  <br> 0.011 0.061 |
| :--- |



Node in 3 S wavefunction near maximum in 1 P wavefunction so large cancellation very sensitive to details of the wavefunctions

$$
\begin{aligned}
& \left\langle 1^{3} P_{2}\right| r\left|3^{3} S_{1}\right\rangle \approx+0.096 \mathrm{GeV}^{-1} \\
& \left\langle 1^{3} P_{1}\right| r\left|3^{3} S_{1}\right\rangle \approx+0.040 \mathrm{GeV}^{-1} \\
& \left\langle 1^{3} P_{0}\right| r\left|3^{3} S_{1}\right\rangle \approx-0.026 \mathrm{GeV}^{-1}
\end{aligned}
$$



Table I: Properties of $\psi(2 S) \rightarrow \gamma \chi_{c J}$ decays, using results from Refs. [54] and [66] as well as Eq. (6).

| $J$ | $k_{\gamma}$ <br> $(\mathrm{MeV})$ | $\mathcal{B}[66]$ <br> $(\%)$ | $\Gamma\left[\psi(2 S) \rightarrow \gamma \chi_{c J}\right]$ <br> $(\mathrm{keV})$ | $\|\langle 1 P\| r\| 2 S\rangle \mid$ <br> $\left(\mathrm{GeV}^{-1}\right)$ |
| :---: | :---: | :---: | :---: | :---: |
| 2 | $127.60 \pm 0.09$ | $9.33 \pm 0.14 \pm 0.61$ | $31.4 \pm 2.4$ | $2.51 \pm 0.10$ |
| 1 | $171.26 \pm 0.07$ | $9.07 \pm 0.11 \pm 0.54$ | $30.6 \pm 2.2$ | $2.05 \pm 0.08$ |
| 0 | $261.35 \pm 0.33$ | $9.22 \pm 0.11 \pm 0.46$ | $31.1 \pm 2.0$ | $1.90 \pm 0.06$ |



Table III: Properties of the transitions $\chi_{c J} \rightarrow \gamma J / \psi$. (Ref. [54]; Eq. (6)).

| $J$ | $k_{\gamma}$ <br> $(\mathrm{MeV})$ | $\Gamma\left(\chi_{c J} \rightarrow \gamma J / \psi\right)$ <br> $(\mathrm{keV})$ | $\langle 1 S\| r\|1 P\rangle \mid$ <br> $(\mathrm{GeV})^{-1}$ |
| :---: | :---: | :---: | :---: |
| 2 | $429.63 \pm 0.08$ | $416 \pm 32$ | $1.91 \pm 0.07$ |
| 1 | $389.36 \pm 0.07$ | $317 \pm 25$ | $1.93 \pm 0.08$ |
| 0 | $303.05 \pm 0.32$ | $135 \pm 15$ | $1.84 \pm 0.10$ |




## Production of the D-wave states

- By direct scans in $e^{+} e^{-}$to produce ${ }^{3} D_{1}\left(J^{P C}=1^{--}\right)$
- Use for $4 \gamma$ E1 cascade to search for $\Upsilon\left(1^{3} D_{j}\right)$

- Estimate the radiative widths and BR using quark model
-In e.m. cascades: $\mathrm{Y}(3 \mathrm{~S}) \rightarrow \gamma \chi_{\mathrm{b}} \rightarrow \gamma \gamma{ }^{3} \mathrm{D}_{\mathrm{J}}$
$\left.\left.\Gamma=\frac{4}{3} e_{Q}^{2} \alpha C\left(J_{i} L_{i} J_{f} L_{f} S\right)|\langle P| r| S\right\rangle\left|\omega^{3} \quad c\left(J_{i} L_{i} J_{J} L_{f} S\right)=\max \left(L_{i}, L_{f}\right)\left(2 J_{f}+1\right)\right| \begin{array}{ll}L_{f} & J_{f} \\ J_{i} & S \\ L_{i} & 1\end{array}\right\}^{2}$
-Some $4 \gamma$ cascades with observable \# of events/ $10^{6} \mathrm{Y}(3 S)$ 's:

| Cascade | Events |
| :--- | :--- |
| $3^{3} \mathrm{~S}_{1} \rightarrow 2^{3} \mathrm{P}_{2} \rightarrow 1^{3} \mathrm{D}_{3} \rightarrow 1^{3} \mathrm{P}_{2} \rightarrow 1^{3} \mathrm{~S}_{1}$ | 7.8 |
| $3^{3} \mathrm{~S}_{1} \rightarrow 2^{3} \mathrm{P}_{2} \rightarrow 1^{3} \mathrm{D}_{2} \rightarrow 1^{3} \mathrm{P}_{1} \rightarrow 1^{3} \mathrm{~S}_{1}$ | 2.7 |
| $3^{3} \mathrm{~S}_{1} \rightarrow 2^{3} \mathrm{P}_{1} \rightarrow 1^{3} \mathrm{D}_{2} \rightarrow 1^{3} \mathrm{P}_{1} \rightarrow 1^{3} \mathrm{~S}_{1}$ | 20 |
| $3^{3} \mathrm{~S}_{1} \rightarrow 2^{3} \mathrm{P}_{1} \rightarrow 1^{3} \mathrm{D}_{1} \rightarrow 1^{3} \mathrm{P}_{1} \rightarrow 1^{3} \mathrm{~S}_{1}$ | 3.3 |

S.G + J. Rosner, Phys Rev D64, 097501 (2001)

Expect ~38 events $/ 10^{6} \mathrm{Y}(3 \mathrm{~S})$ via ${ }^{3} \mathrm{D}_{\mathrm{J}}$
-The $\mathrm{e}^{+} \mathrm{e}^{-}$final states leads to less background

- $\mu^{+} \mu^{-}$final states also contribute if $\mu^{\prime}$ s are identified


## CLEO finds:

$\mathrm{B}\left(\Upsilon(3 \mathrm{~S}) \mapsto \gamma \mathrm{Y}^{\prime}(1 \mathrm{D}) \mapsto \gamma \gamma \gamma \mathrm{Y}(1 \mathrm{~S}) \mapsto \gamma \gamma^{\left.\ell^{++} \ell^{-}\right)}=(3.3 \pm 0.6 \pm 0.5) 10^{-5}\right.$ (vs GR prediction of $3.8 \times 10^{-5}$ )


- Mass averaged over different fits: $10162.2 \pm 1.6 \mathrm{MeV}$
- Inconsistent with the $\mathrm{r}\left(1 \mathrm{D}_{3}\right)$
- Could be the $r\left(1 D_{2}\right)$ or $r\left(1 D_{1}\right)$
- The theory predicts the rate ratio: $\Upsilon\left(1 D_{2}\right) / \Upsilon\left(1 D_{1}\right)=6$
- Thus, the $\mathrm{r}\left(1 \mathrm{D}_{2}\right)$ is the most likely interpretation


## M1 Transitions

Because quarks have spin they may emit a photon via a spin flip - The magnetic dipole transition

To obtain the interaction Hamiltonian we perform a non-relativistic reduction of

$$
\begin{aligned}
& H_{I}=e \int d x j_{e m}^{\mu}(x) A_{\mu}(x) \\
& \text { where } j_{e m}^{\mu}(x)=\bar{q}(x) Q \gamma^{\mu} q(x)
\end{aligned}
$$

We expand the Dirac spinors to lowest order in $\mathrm{p} / \mathrm{m}$ Denoting the large and small components by $q_{1}$ and $q_{2}$

$$
q_{2}(x)=-\frac{i \vec{\sigma} \cdot \vec{\nabla}}{2 m} q_{1}(x)
$$

$$
\vec{j}_{e m}(x)=\frac{-i}{2 m}\left[q_{1}^{\dagger} Q\left(\nabla q_{1}\right)-\left(\nabla q_{1}^{\dagger}\right) Q q_{1}+i \nabla \times q_{1}^{\dagger} Q \vec{\sigma} q_{1}\right]
$$

So the interaction Hamiltonian is given by:

$$
H_{I}=\frac{-e Q}{2 m}[\vec{A}(\vec{r}) \cdot \vec{p}+\vec{p} \cdot \vec{A}(\vec{r})+\vec{\sigma} \cdot[\vec{\nabla} \times \vec{A}(\vec{r})]
$$

So:

$$
\begin{aligned}
\langle 0| H_{I}|\gamma(\vec{k}, \epsilon)\rangle= & -\frac{1}{(2 \pi)^{3 / 2}} \frac{1}{(2 \omega)^{1 / 2}} e Q \frac{1}{2 m}\left[e^{i \vec{k} \cdot \vec{r}} \vec{\epsilon} \cdot \vec{p}+\right. \\
& \left.\vec{\epsilon} \cdot \vec{p} e^{i \vec{k} \cdot \vec{r}}+i \vec{\sigma} \cdot(\vec{k} \times \vec{\epsilon}) e^{i \vec{k} \cdot \vec{r}}\right]
\end{aligned}
$$

(For antiquarks change the sign of the charge)
$\mu=\frac{e}{2 m_{1}} \quad$ Is the magnetic dipole moment of the quark
For magnetic dipole transitions:

$$
\begin{aligned}
& M_{i f}=i \mu\langle f| \vec{\sigma}|i\rangle \cdot \vec{k} \times \vec{\epsilon}^{*} \\
& \quad \vec{\epsilon}=\frac{1}{\sqrt{2}}(1, \pm i, 0) \\
& \left|\begin{array}{ccc}
\sigma_{x} & \sigma_{y} & \sigma_{z} \\
k_{x} & k_{y} & k_{z} \\
1 & i & 0
\end{array}\right|=i \sigma_{z}\left(k_{x}+i k_{y}\right)-i k_{z} \sigma_{x}+k_{z} \sigma_{y}
\end{aligned}
$$

Choosing $z$ as the $\gamma$ direction

$$
M_{i f}=-\frac{i e_{q}}{2 m} k_{\gamma}\langle f| \sigma_{x}-i \sigma_{y}|i\rangle \text { where } \sigma_{x}-i \sigma_{y}=\sigma_{-}
$$

$$
\text { if instead take } \vec{k}=k_{y}
$$

$$
\begin{aligned}
M_{i f} & =-\frac{i e_{q}}{2 m} k_{\gamma}\langle f| \sigma_{z}|i\rangle \\
& =k_{\gamma} \sqrt{2 M_{i}} \sqrt{2 M_{f}} \int d^{3} r \psi_{f}^{*}(r) \psi_{i}(r) \times\langle f| \sum \mu_{i} \sigma_{z i}|i\rangle
\end{aligned}
$$

$$
\begin{aligned}
& \text { e.g. } J / \psi \rightarrow \eta_{c} \gamma\left({ }^{3} S_{1} \rightarrow{ }^{1} S_{0} \gamma\right) \\
& \begin{aligned}
& A\left({ }^{3} S_{1} \rightarrow{ }^{1} S_{0} \gamma\right)=- \\
& i k_{\gamma} \sqrt{2 M_{i}} \sqrt{2 M_{f}}\langle f \mid i\rangle \\
& \left.\times\left\langle\sqrt{\frac{1}{2}}(\uparrow \downarrow-\downarrow \uparrow)\right| \frac{e_{q}}{2 m_{q}} \frac{\left(\sigma_{x}-i \sigma_{y}\right)_{q}}{\sqrt{2}}+\mu_{\bar{q}} \frac{\left(\sigma_{x}-i \sigma_{y}\right)_{\bar{q}}}{\sqrt{2}} \right\rvert\, \\
&=-i k_{\gamma} \sqrt{2 M_{i}} \sqrt{2 M_{f}}\langle f \mid i\rangle\left[\frac{-e_{q}}{2 m_{q}}+\frac{e_{\bar{q}}}{2 m_{\bar{q}}}\right] \\
&=-i k_{\gamma} \sqrt{2 M_{i}} \sqrt{2 M_{f}}\langle f \mid i\rangle \frac{e e_{q}}{m_{c}}
\end{aligned} \\
& \Rightarrow \frac{d \Gamma}{d \Omega}=k_{\gamma} \frac{4 \pi \alpha}{8 \pi^{2}} k_{\gamma}^{2}|\langle f \mid i\rangle|^{2} \frac{e_{c}^{2}}{m_{c}^{2}} \\
& \quad \text { averaging over angles gives the total width }
\end{aligned}
$$

$$
\Gamma=\frac{k_{\gamma}^{3}}{3 \pi}|\langle f \mid i\rangle|^{2} \frac{e_{c}^{2}}{m_{c}^{2}}
$$

$$
\text { Take }\langle f \mid i\rangle=1 \quad \omega=115
$$

$$
\text { so } \Gamma=0.19 \mathrm{MeV} \text { vs } 0.88 \mathrm{keV} \text { (expt) }
$$

What about? $2^{3} S_{1} \rightarrow 1^{1} S_{0}$

$$
\langle f \mid i\rangle=0 \text { since } 2 S \perp 1 S
$$

The decay $\psi(2 S) \rightarrow \gamma \eta_{c}(1 S)$ is a forbidden magnetic dipole (M1) transition
The photon energy is 638 MeV , leading to a non-zero matrix element $\langle 1 S| j_{0}(k r / 2)|2 S\rangle$.

$$
\begin{gathered}
\Gamma\left[\psi(2 S) \rightarrow \gamma \eta_{c}(1 S)=(1.00 \pm 0.16)\right. \\
\left.\left|\langle 1 S| j_{0}(k r / 2)\right| 2 S\right\rangle \mid=0.045 \pm 0.004
\end{gathered}
$$

The $\left.h_{c}\left(1^{1} P_{1}\right) \quad \Gamma\left[h_{b}\left(1^{1} P_{1}\right) \rightarrow \eta_{b}\left(1^{1} S_{0}\right)+\gamma\right]=\frac{4}{9} \alpha e_{Q}^{2}\left|\left\langle{ }^{1} S_{0}\right| r\right|^{1} P_{1}\right\rangle\left.\right|^{2} \omega^{3}=37 \mathrm{keV}$


## M1 transitions: production of $\eta_{b}(n S)$ states

S.G + J. Rosner, Phys Rev D64, 074011 (2001)

Proceeds via magnetic dipole (M1) transitions:

$$
\begin{gathered}
\mathrm{Y}(\mathrm{nS}) \rightarrow \eta\left(\mathrm{n}^{\prime} \mathrm{S}\right)+\gamma \\
\left.\Gamma{ }^{3} \mathrm{~S}_{1} \rightarrow S^{1} \mathrm{~S}_{0}+\gamma\right)=\left.\frac{4}{3} \alpha \frac{e_{Q}^{2}}{m_{Q}^{2}}\left\langle\langle f| j_{0}(k r / 2) \mid i\right\rangle\right|^{2} \omega^{z}
\end{gathered}
$$



- Hindered transitions have large phase space
-Relativistic corrections resulting in differences in ${ }^{3} S_{1}$ and ${ }^{1} S_{0}$ wavefunctions due to hyperfine interaction

|  | Transition | BR $\left(10^{-4}\right)$ |
| :--- | :--- | :--- |
| $\mathrm{Y}(3 \mathrm{~S})$ |  |  |
| $\left(\Gamma_{\text {tot }}=52.5 \mathrm{keV}\right)$ | $\rightarrow 3^{1} S_{0}$ | 0.10 |
|  | $\rightarrow 2^{1} S_{0}$ | 4.7 |
|  | $\rightarrow 1^{1} S_{0}$ | 25 |
| $\mathrm{Y}(2 \mathrm{~S})$ | $\rightarrow 2^{1} S_{0}$ | 0.21 |
| $\left(\Gamma_{\mathrm{tot}}=44 \mathrm{keV}\right)$ | $\rightarrow 1^{1} S_{0}$ | 13 |
| $\mathrm{Y}(1 \mathrm{~S})$ | $\rightarrow 1^{1} S_{0}$ | 2.2 |
| $\left(\Gamma_{\mathrm{tot}}=26.3 \mathrm{keV}\right)$ |  |  |

- Expect substantial rate to produce $\eta_{b}$ 's
- Also $\mathrm{Y}(3 \mathrm{~S}) \rightarrow h_{b}\left({ }^{1} \mathrm{P}_{1}\right) \pi \pi \rightarrow \eta_{\mathrm{b}}+\gamma+\pi \pi$

$$
B R=0.1-1 \% \quad B R=50 \%
$$

[Kuang \& Yan PRD24, 2874 (1981); Voloshin Yad Fiz 43, 1571 (1986)]

## Decays:

$$
\begin{aligned}
& J / \psi \rightarrow e^{+} e^{-} \\
& \left({ }^{3} S_{1} \rightarrow e^{+} e^{-}\right) \\
& \stackrel{c}{c}>\sim^{r}<l^{l} \\
& A\left(V_{i} \rightarrow e^{+} e^{-}\right) \equiv\left\langle e^{+} e^{-}\right| M\left|V_{i}\right\rangle \\
& =\frac{4 \pi \alpha e_{q}}{M^{2}}\left\langle e^{+} e^{-}\right| j_{k}^{(e m)}|0\rangle\langle 0| j_{k}\left|V_{i}\right\rangle \\
& =\frac{4 \pi \alpha e_{q}}{M^{2}} \bar{U}_{e}\left(-p_{+}\right) \gamma_{k} U\left(p_{-}\right)\langle 0| j_{k}\left|V_{i}\right\rangle \\
& \langle 0| j_{k}\left|V_{i}\right\rangle=\sqrt{3 \times 2 M} \int d^{3} p \phi_{s}(p) Y_{00}\langle 0| j_{e m}^{\mu}|c \bar{c}\rangle \\
& \text { where } \sum_{\text {colour }} \sqrt{\frac{1}{3}}(r \bar{r}+b \bar{b}+g \bar{g})=\frac{3}{\sqrt{3}}=\sqrt{3}
\end{aligned}
$$

Typically express the matrix element in the form:

$$
\langle 0| j_{e m}^{\mu}(0)|\psi(k, \lambda)\rangle=\frac{\epsilon^{\mu}(k, \lambda)}{(2 \pi)^{3 / 2}} f_{\psi}
$$

## In non-relativistic limit

$$
\begin{aligned}
\Rightarrow\langle 0| j_{e m}^{\mu}(0)|V(\uparrow)\rangle & =\sqrt{12 M} \epsilon^{\mu}(\uparrow) \psi_{S}(0) \\
& \equiv \epsilon^{\mu}(k, \lambda) f_{V} \\
f_{V} & =\sqrt{12 M} \psi_{S}(0)
\end{aligned}
$$

$$
\begin{aligned}
\Gamma & =\frac{1}{2 M} \int|M|^{2} \frac{m_{e}}{E_{e^{+}}} \frac{m_{e}}{E_{e^{-}}} \frac{d^{3} p_{+}}{(2 \pi)^{3}} \frac{d^{3} p_{-}}{(2 \pi)^{3}}(2 \pi)^{4} \delta^{4}\left(P-p_{+}-p_{-}\right) \\
& =\frac{e_{Q}^{2} e^{4}}{12 \pi M^{3}}\left(12 M(2 \pi)^{3}|\psi(0)|^{2}\right) \\
& =\frac{16 \pi^{2} \alpha^{2} e_{Q}^{2}}{\pi M^{3}} M|\psi(0)|^{2} \\
& =\frac{16 \pi \alpha^{2} e_{Q}^{2}}{M^{2}}|\psi(0)|^{2} \\
& \psi_{S}(0)=\frac{1}{\sqrt{4 \pi}} R(0)
\end{aligned}
$$

## What about $\psi^{\prime \prime}(3770) ? \quad e^{+} e^{-} \rightarrow \psi^{\prime \prime}(3770)$

${ }^{3} D_{1}$ state so expect $\Gamma=0$ since $\psi_{D}(0)=0$ but not so

$$
\begin{aligned}
|V(\uparrow)\rangle= & \sqrt{6 M} \int d^{3} p \phi_{D}(p)\left\{\sqrt{3 / 5} Y_{2+2}(\theta, \phi)|q(\downarrow) \bar{q}(\downarrow)\rangle\right. \\
& \left.-\sqrt{3 / 10} Y_{2+1}(\theta, \phi)|q(\uparrow) \bar{q}(\downarrow)\rangle+\sqrt{1 / 10} Y_{20}(\theta, \phi)|q(\uparrow) \bar{q}(\uparrow)\rangle\right\}
\end{aligned}
$$

After much work get:

$$
\begin{aligned}
\langle 0| j_{e m}^{\mu}(0)|V(\uparrow)\rangle & =\frac{\sqrt{12 M}}{(2 \pi)^{3}} \epsilon^{\mu}(\uparrow) \int d^{3} p \frac{\phi_{D}(p)}{\sqrt{32 \pi}} \frac{4}{3} \frac{p^{2}}{E(E+m)} \\
\lim _{x \rightarrow 0} \int d^{3} p \phi_{D}(p) \frac{p^{2}}{2 m^{2}} \frac{e^{i \vec{p} \cdot \vec{x}}}{(2 \pi)^{3 / 2}} & =-\frac{1}{2 m^{2}} \lim _{x \rightarrow 0} \frac{\partial^{2}}{\partial x_{i}^{2}} \int d^{3} p \frac{e^{i \vec{p} \cdot \vec{x}}}{(2 \pi)^{3 / 2}} \phi_{D}(p) \\
& =-\frac{1}{2 m^{2}} \frac{\partial^{2} R_{D}(0)}{\partial r^{2}}=-\frac{1}{2 m^{2}} R_{D}^{\prime \prime}(0)
\end{aligned}
$$

In general, for state of angular momentum $L$ get $R^{(L)}(0)$

More carefully get:

$$
\begin{aligned}
& \langle 0| j_{e m}^{\mu}(0)|V(\uparrow)\rangle=\frac{\sqrt{12 M}}{(2 \pi)^{3 / 2}} \frac{5}{4} \frac{R^{\prime \prime}(0)}{m^{2} \sqrt{2 \pi}} \epsilon^{\mu}(\uparrow) \\
& \Gamma=\frac{\alpha^{2}\left(e_{q} / e\right)^{2}}{M_{V}^{2}} \frac{25}{2} \frac{\left|R_{D}^{\prime \prime}(0)\right|^{2}}{m_{q}^{2}}
\end{aligned}
$$

## Also have decays to hadronic final states:

$$
\begin{aligned}
n_{c} & \rightarrow \gamma \gamma \\
& \rightarrow y_{j} \\
J / E & \rightarrow \gamma \gamma \gamma \\
& \rightarrow \gamma y y \\
& \rightarrow g_{j \gamma}
\end{aligned}
$$



$$
d t .
$$



Start with annhilation rates for positronium:

$$
\begin{aligned}
& \Gamma\left({ }^{1} S_{0} \rightarrow 2 \gamma\right)=\frac{4 \pi \alpha^{2}}{m^{2}}\left|\psi_{S}(0)\right|^{2}=\frac{4 \alpha^{2}}{m^{2}}\left|R_{S}(0)\right|^{2} \\
& \Gamma\left({ }^{3} P_{0} \rightarrow 2 \gamma\right)=\frac{256}{3} \frac{\alpha^{2}}{m^{4}}\left|R_{P}^{\prime}(0)\right|^{2} \\
& \Gamma\left({ }^{3} P_{2} \rightarrow 2 \gamma\right)=\frac{4}{15} \Gamma\left({ }^{3} P_{0} \rightarrow \gamma \gamma\right)\left(\frac{M_{0}}{M_{2}}\right) \\
& \Gamma\left({ }^{3} S_{1} \rightarrow 3 \gamma\right)=\frac{16}{9 \pi}\left(\pi^{2}-9\right) \frac{\alpha^{3}}{m^{2}}\left|R_{S}(0)\right|^{2}
\end{aligned}
$$

To relate to hadron decays include quark charges For decays to gluons must include $\alpha_{S}$ and $\lambda$ 's for each gluon

$$
\begin{aligned}
& =\frac{\alpha_{s}}{e_{q}^{2} \alpha} \frac{\left(\lambda_{a} / 2\right)_{j}^{i}\left(\lambda_{b} / 2\right)_{i}^{j}}{\delta_{j}^{i} \delta_{i}^{j}}=\frac{\alpha_{s}}{\alpha e_{q}^{2}} \frac{\operatorname{Tr}\left(\lambda_{a} / 2 \lambda_{b} / 2\right)}{3}=\frac{\alpha_{s}}{\alpha} \frac{\frac{1}{2} \delta_{a b}}{e_{q}^{2} 3} \\
& \text { where } \operatorname{Tr}\left(\lambda_{a} / 2 \lambda_{b} / 2\right)=\frac{1}{2} \delta_{a b} \\
& \Rightarrow \frac{\Gamma(2 g)}{\Gamma(2 \gamma)}=\frac{2}{9} \frac{\alpha_{s}^{2}}{\alpha^{2} e_{q}^{4}}
\end{aligned}
$$

For 3 gluons/photons:

$$
\begin{aligned}
& \frac{M(3 g)}{M(3 \gamma)}=\frac{\alpha_{s}^{3 / 2}}{e_{q}^{3} \alpha^{3 / 2}} \frac{\left(\lambda_{a} / 2\right)_{j}^{i}\left(\lambda_{b} / 2\right)_{k}^{j}\left(\lambda_{c} / 2\right)_{i}^{k}}{\delta_{j}^{i} \delta_{k}^{j} ; \delta_{i}^{k}}=\frac{\alpha_{s}^{3 / 2}}{e_{q}^{3} \alpha^{3 / 2}} \frac{1}{2} \frac{\left.\operatorname{Tr}\left(\left\{\lambda_{a} / 2, \lambda_{b} / 2\right)\right\} \lambda_{c} / 2\right)}{\delta_{j}^{i} \delta_{k}^{j} \delta_{i}^{k}} \\
& \Rightarrow \frac{\Gamma(2 g)}{\Gamma(2 \gamma)}=\frac{5}{54} \frac{\alpha_{s}^{3}}{\alpha^{3} e_{q}^{6}} \text { where } \sum_{a, b, c}\left(d_{a b c}\right)^{2}=40 / 3
\end{aligned}
$$

$$
\begin{aligned}
& \Gamma\left(\eta_{c} \rightarrow 2 \gamma\right)=12 \alpha^{2} e_{q}^{4} \frac{\left|R_{S}(0)\right|^{2}}{M^{2}} \\
& \Gamma\left(\eta_{c} \rightarrow 2 g\right)=\frac{8}{3} \alpha_{S} \frac{\left|R_{S}(0)\right|^{2}}{M^{2}} \\
& \Gamma(J / \psi \rightarrow 3 \gamma)=\frac{16\left(\pi^{2}-9\right) \alpha^{3}}{3} e_{q}^{6} \frac{\left|R_{S}(0)\right|^{2}}{M^{2}} \\
& \Gamma(J / \psi \rightarrow 3 g)=\frac{40}{81 \pi}\left(\pi^{2}-9\right) \alpha_{s}^{3} \frac{\left|R_{S}(0)\right|^{2}}{M^{2}} \\
& \Gamma(J / \psi \rightarrow 2 g \gamma)=\frac{32}{9 \pi}\left(\pi^{2}-9\right) \alpha_{s}^{2} \alpha e_{1}^{2} \frac{\left|R_{S}(0)\right|^{2}}{M^{2}}
\end{aligned}
$$

## For Completeness:

$$
\begin{aligned}
& \Gamma\left(\chi_{0} \rightarrow 2 g\right)=96 \alpha_{s}^{2} \frac{\left|R_{\chi_{o}}^{\prime}(0)\right|^{2}}{M_{\chi_{0}}^{4}} \\
& \Gamma\left(\chi_{1} \rightarrow q \bar{q} g\right)=\frac{n_{f}}{3} \frac{128}{3 \pi} \alpha_{s}^{3} \frac{\left|R_{\chi_{1}}^{\prime}(0)\right|^{2}}{M_{\chi_{0}}^{4}} \ln \left(\frac{4 m_{c}^{2}}{4 m_{c}^{2}-M_{\chi}^{2}}\right) \\
& \Gamma\left(\chi_{2} \rightarrow 2 g\right)=\frac{128}{5} \alpha_{s}^{2} \frac{\left|R_{\chi_{2}}^{\prime}(0)\right|^{2}}{M_{\chi_{0}}^{4}} \\
& \Gamma\left(h_{c} \rightarrow q \bar{q} g\right)=\frac{320}{9 \pi} \alpha_{s}^{3} \frac{\left|R_{h_{c}}^{\prime}(0)\right|^{2}}{M_{h_{c}}^{4}} \ln \left(\frac{4 m_{c}^{2}}{4 m_{c}^{2}-M_{h_{c}}^{2}}\right)
\end{aligned}
$$

## 4. What about mesons with light quarks?

Historically, it was the successes of the quark model that led many physicists to believe that the quark model has something to do with reality


Essential features are the same, except:
-Relative importance of relativistic effects
-Hyperfine splittings are comparable in size to orbital splittings
Conclude
-potential models approximately valid

## Hadron masses in a gauge theory*

A. De Rújula, Howard Georgi, ${ }^{\dagger}$ and S. L. Glashow

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(Received 24 February 1975)
We explore the implications for hadron spectroscopy of the "standard" gauge model of weak, electromagnetic, and strong interactions. The model involves four types of fractionally charged quarks, each in three colors, coupling to massless gauge gluons. The quarks are confined within colorless hadrons by a long-range spinindependent force realizing infrared slavery. We use the asymptotic freedom of the model to argue that for the calculation of hadron masses, the short-range quark-quark interaction may be taken to be Coulomb-like. We rederive many successful quark-model mass relations for the low-lying hadrons. Because a specific interaction and symmetry-breaking mechanism are forced on us by the underlying renormalizable gauge field theory, we also obtain new mass relations. They are well satisfied. We develop a qualitative understanding of many features of the hadron mass spectrum, such as the origin and sign of the $\boldsymbol{\Sigma}-\Lambda$ mass splitting. Interpreting the newly discovered narrow boson resonances as states of charmonium, we use the model to predict the masses of charmed mesons and baryons.

Flavour content:

$$
\begin{aligned}
& \left|\rho^{+}\right\rangle,\left|\pi^{+}\right\rangle=-|u \bar{d}\rangle \\
& \left|\rho^{0}\right\rangle,\left|\pi^{0}\right\rangle=\frac{1}{\sqrt{2}}|u \bar{u}-d \bar{d}\rangle \\
& |\omega\rangle=\frac{1}{\sqrt{2}}|u \bar{u}+d \bar{d}\rangle \\
& |\eta\rangle=\frac{1}{\sqrt{6}}|u \bar{u}+d \bar{d}-2 s \bar{s}\rangle \\
& \left|\eta^{\prime}\right\rangle=\frac{1}{\sqrt{3}}|u \bar{u}+d \bar{d}+s \bar{s}\rangle \\
& |\phi\rangle=|s \bar{s}\rangle \\
& \left|K^{+}\right\rangle=|u \bar{s}\rangle \\
& \left|K^{0}\right\rangle=|d \bar{s}\rangle \\
& \left|\bar{K}^{0}\right\rangle=-|s \bar{d}\rangle \\
& \left|K^{-}\right\rangle=|s \bar{u}\rangle
\end{aligned}
$$

## In heavy quarkonium we used:

$$
\begin{aligned}
& M=m_{1}+m_{2}+E_{n l} \\
& {\left[\frac{p^{2}}{2 \mu}+V(r)\right] \psi=E_{n l} \psi}
\end{aligned}
$$

$$
H_{i j}^{c o n f}=-\frac{4}{3} \frac{\alpha_{s}(r)}{r}+b r
$$

This is a non-relativistic formula ( $\mathrm{v} / \mathrm{c}$ ) $=\quad b \bar{b} \quad 0.26$
$\begin{array}{ll}c \bar{C} & 0.45\end{array}$
$\begin{array}{ll}s \bar{s} & 0.78\end{array}$
What do we do?
$\begin{array}{ll}u \bar{u} & 0.9\end{array}$

- Use it anyway and see what happens. Taking this approach the general features are OK
- Try to relativize it.


## Spin dependent interactions:

$$
\Delta\left[M\left({ }^{3} S_{1}\right)-M\left({ }^{1} S_{0}\right)\right]=\frac{3 \pi \alpha_{s}}{9 m_{1} m_{2}}|\psi(0)|^{2}
$$

Approximate ${ }^{3} S_{1}$ and ${ }^{1} S_{0}$ masses by:

$$
\begin{aligned}
& M\left({ }^{3} S_{1}\right)=M(S)+\frac{1}{4} \frac{a}{m_{q} m_{\bar{q}}} \\
& M\left({ }^{1} S_{0}\right)=M(S)-\frac{3}{4} \frac{a}{m_{q} m_{\bar{q}}}
\end{aligned}
$$

If a is approximately constant:

$$
\begin{aligned}
& \frac{M(\rho)-M(\pi)}{M\left(K^{*}\right)-M(K)} \approx \frac{m_{u} m_{s}}{m_{u} m_{u}} \approx \frac{m_{s}}{m_{u}} \approx \frac{500}{300} \approx 1.7 \\
& \frac{770-140}{892-495} \approx \frac{630}{400} \approx 1.7
\end{aligned}
$$

Similarly:

$$
\begin{aligned}
& \frac{M\left(K^{*}\right)-M(K)}{M\left(D^{*}\right)-M(D)} \approx \frac{m_{u} m_{c}}{m_{u} m_{s}} \approx \frac{m_{c}}{m_{s}} \approx \frac{1.6}{0.55} \approx 2.9 \\
& \frac{892-494}{2010-1870} \approx \frac{400}{140} \approx 2.9
\end{aligned}
$$

## So splittings reasonably well described

Because ${ }^{3} P_{c o g}-{ }^{-1} P_{1}$ splitting is small supports short range contact interaction

## Electromagnetic transitions:

As before: $\quad \Gamma_{M 1}=\left.\frac{k_{r}^{3}}{3 \pi}\left\langle\left.\langle\mid i\rangle\right|^{2}\right| \sum \mu_{i} \sigma_{z i}\right|^{2}$
For example:

$$
\begin{aligned}
& K^{*+} \rightarrow K^{+} \gamma \\
& \left\langle u \bar{s} \frac{1}{\sqrt{2}}(\uparrow \downarrow-\downarrow \uparrow)\right| \frac{e_{i}}{2 m_{i}} \sigma_{\bar{z}}\left|u \bar{s} \frac{1}{\sqrt{2}}(\uparrow \downarrow-\downarrow \uparrow)\right\rangle \\
& =\frac{1}{2}\langle u \bar{s}| \frac{e_{q}}{2 m_{q}}+\frac{e_{q}}{2 m_{q}}-\frac{e_{\bar{q}}}{2 m_{\bar{q}}}-\frac{e_{\bar{q}}}{2 m_{\bar{q}}}|u \bar{s}\rangle \\
& =\frac{1}{2}\left[\frac{e_{u}}{m_{u}}-\frac{e_{s}}{m_{s}}\right]=\frac{1}{2}\left[\frac{2}{3} \frac{1}{m_{u}}-\frac{1}{3} \frac{1}{m_{s}}\right]
\end{aligned}
$$

## Strong (Zweig allowed) Decays:

A number of models to calculate strong decays.
Give good qualitative agreement with experiment with only 1 free parameter (using QM wavefunctions)

Important input to disentangle hadron spectrum

## Relativistic effects:

Clearly light quark hadrons are relativistic
Various attempts to "relativize" QM
Generally improves agreement
But much is missing. Major battles about what is correct approach.

BUT QM seems to get the physics right.
"Better to get the right degrees of freedom"



- Many unconfirmed states: $\mathrm{f}_{1}$ (1530), $\mathrm{h}_{1}$ (1380)
- Many puzzles:
$\eta(1440), f_{1}(1420), f_{0}(1500) f_{J}(1710), f_{J}(2200)$
S. Godfrey, Carleton University

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## Scientists find mystery particle

By Dr David Whitehouse
BBC News Online science editor

## Scientists have found a sub-atomic particle they cannot explain using current theories of energy and matter. <br> The discovery was made by researchers based at the High Energy Accelerator Research Organisation in Tsukuba.



Fermilab confirmed the discovery
Classified as $\mathrm{X}(3872)$, the particle was seen fleetingly in an atom smasher and has been dubbed the "mystery meson".

The Japanese team says understanding its existence may require a change to the Standard Model, the accepted theory of the way the Universe is constructed.

## An eternity

$X$ (3872) was found among the decay products of so-called beauty mesons - sub-atomic particles that are produced in large numbers at the Tsukuba "meson factory".

It weighs about the same as a single atom of helium and exists for only about one billionth of a trillionth of a second before it decays into other longer-lived, more familiar

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## $X(3943), Y(3943)$, and $Z(3931)$



## $2 P$ or not $2 P$ that is the question!

S. Godfrey, Carleton University

## X(3940)

Seen by Belle recoiling against $J / \psi$ in $e^{+} e^{-}$collisions

$$
\begin{aligned}
& M=3943 \pm 6 \pm 6 \mathrm{MeV} \\
& \Gamma<52 \mathrm{MeV}
\end{aligned}
$$

$$
B R\left(X \rightarrow D D^{*}\right)=96^{+45}-32 \pm 22 \%
$$

$$
B R(X \rightarrow D D)<41 \% \quad(90 \% C L)
$$



Suggests unnatural parity state
$\mathrm{BR}(\mathrm{X} \rightarrow \omega \mathrm{J} / \psi)<26 \%(90 \% \mathrm{CL})$

- Decay to DD* but not DD suggests unnatural parity state
- Belle speculates that $X$ is $3^{1} S_{0}$ given the $3^{3} S_{1} \psi(4040)$ - Mass is roughly correct
-. $\eta_{\mathrm{c}}$ and $\eta_{\mathrm{c}}^{\prime}$ are also produced in double charm production
See also Eichten Lane Quigg PRD73 014014(2006)
-Predicted width for $3^{1} \mathrm{~S}_{0}$ with $M=3943 \sim 50 \mathrm{MeV}$ close to $\Gamma(X(3943))$ upper bound
- Identification of $\psi(4040)$ as $3^{3} S_{1}$ state implies hyperfine splitting 88 MeV with X(3943)
-Larger than the 25 splitting and larger than predicted in potential models
- Discrepancy could be due to:
- Difficulty in fitting true pole position of $3^{3} S_{1}$ state
- Nearby thresholds with s-wave + p-wave charm mesons so possibly stronger threshold effects
What of ${ }^{1}{ }^{1}$ S. Godrey, Carleon $\eta_{\text {Giversity }}$ assignment is search for this state in $\gamma \gamma \underset{83}{\rightarrow} \mathrm{DD}^{*}$


## Y(3940)

See in $\omega \mathrm{J} / \psi$ subsystem of the decay $\mathrm{B} \rightarrow \mathrm{K} \pi \pi \pi \mathrm{J} / \psi$
Belle: Phys. Rev. Lett. 94, 182002 (2005)
$M=3943 \pm 11 \pm 13 \mathrm{MeV}$
$\Gamma=87 \pm 22 \pm 26 \mathrm{MeV}$
Not seen in $Y \rightarrow$ DD or DD*

Mass and width suggest radially excited P -wave charmonium

But $\omega \mathrm{J} / \psi$ decay mode is peculiar:
 $\mathrm{BR}(\mathrm{B} \rightarrow \mathrm{KY}) \mathrm{BR}(\mathrm{Y} \rightarrow \omega \mathrm{J} / \psi)=7.1 \pm 1.3 \pm 3.1 \bullet 10^{-5}$ where one expects $B R\left(B \rightarrow K^{\prime}{ }_{c J}^{\prime}\right)<B R\left(B \rightarrow K \chi_{c J}\right)=4 \bullet 10^{-4}$

Implies $\mathrm{BR}(\mathrm{Y} \rightarrow \omega \mathrm{J} / \psi)>12 \%$ which is unusual for state above open charm threshold

Possibility is $2^{3} P_{1}$ cc state: identifyies $Y(3943)$ as $2 P \chi_{c 1}^{\prime}$
-DD* is the dominant decay mode

- Width consistent with Y(3943): $\Gamma=135 \mathrm{MeV}$
-. $\chi_{c 1}$ is seen in $B$ decays
$\cdot 1^{++} \rightarrow \omega \mathrm{J} / \psi$ is unusual
-but corresponding $\chi_{b 1,2}^{\prime} \rightarrow \omega \mathrm{Y}(1 \mathrm{~S})$ also seen
- Maybe rescattering: $1^{++} \rightarrow D^{*} \rightarrow \omega \mathrm{~J} / \psi$
- Maybe due to mixing with $1^{++}$molecular state $X(3872)$ ?
-Important to - look for DD and DD*
- study angular distributions to DD and DD*
- Observed by Belle in $\gamma \gamma \rightarrow$ DD

$$
\begin{aligned}
& M=3929 \pm 5 \pm 2 \mathrm{MeV} \\
& \Gamma=29 \pm 10 \pm 2 \mathrm{MeV}
\end{aligned}
$$

- Two photon width:

$$
\Gamma_{r v} \bullet \mathrm{~B}_{\mathrm{DD}}=0.18 \pm 0.05 \pm 0.03 \mathrm{keV}
$$

-DD angular distribution consistent with J=2
-Below D* D* threshold
S. Godfrey, Carleton University



- Obvious candidate for $\chi_{c 2}^{\prime}$ (the $\chi_{c 1}^{\prime}$ cannot decay to DD)
- Predicted $\chi_{c 2}^{\prime}$ mass is 3972
$\Gamma\left(\chi_{c 2}^{\prime} \rightarrow \mathrm{DD}\right)=21.5 \mathrm{MeV}$
$\Gamma\left(\chi_{c 2}^{\prime} \rightarrow \mathrm{DD}^{*}\right)=7.1 \mathrm{MeV}$
$\Gamma=47 \mathrm{MeV}$ assuming $M\left(\chi^{\prime}{ }_{c 2}\right)=3931$
- In reasonable agreement with experiment
- Predicted $\mathrm{BR}\left(\chi_{c 2}^{\prime} \rightarrow \mathbf{D D}\right)=70 \% \Rightarrow \Gamma_{\gamma \gamma}{ }^{*} \mathrm{~B}_{\mathrm{DD}}=0.47 \mathrm{keV}$ ( $\mathrm{r}_{y}$ from T.Barnes, IX ${ }^{\text {th }}$ Int.|.Conf. on $\gamma y$ Collisions, La Jolla, 1992.)
- Observed two-photon width about $1 / 2$ predicted value for $\chi_{c 2}^{\prime}$

Could further study $2^{3} \mathrm{P}_{\mathrm{J}}$ states via radiative transitions:
Can find all three ${ }^{3} 2 \mathrm{P}_{\mathrm{J}}$ cc states using $\psi(4040)$ and $\psi(4160) \rightarrow \gamma \mathrm{DD}, \gamma \mathrm{DD}^{*}$

All three E1 rad BFs of the $\psi(4040)$ are $\sim 0.5 * 10^{-3}$.
These would further test whether the $Z, X, Y$ (3.9) are 2P cc


X(3872)
New state $1^{\text {st }}$ observed by Belle: X(3871) Phys Rev. Lett. 91, 2622001 (2003) [hep-ex/0309032]
Confirmed by: CDF Phys Rev. Lett. 93, 072001 (2004)
DO Phys Rev. Lett. 93, 162002 (2004)
BABAR Phys Rev. D71, 071103 (2005)

$$
\begin{aligned}
& M=3872.0 \pm 0.6 \pm 0.5 \mathrm{MeV} \quad \Gamma<2.3 \mathrm{MeV} \text { at 90\% C.L. } \\
& \text { width consistent with detector resolution. }
\end{aligned}
$$

\author{

1. $D^{0} D^{* 0}$ molecule <br> 2. A charmonium hybrid <br> 3. $2^{3} \mathrm{P}_{\mathrm{J}} 1^{3} \mathrm{D}_{2}$ state? <br> 4. Glueball?
}

## Charmonium Options for the $X(3872)$

T.Barnes,S.Godfrey, Phys Rev D69, 050400 (2004) [hep-ph/0311162]

Eichten, Lane \& Quigg, Phys Rev D69, 094019 (2004) [hep-ph/0401210]
Barnes, Godfrey \& Swanson, in preparation
New state $1^{1 s t}$ observed by Belle: X(3871) hep-ex/0309032
Observation of a new narrow charmonium state in exclusive

$$
B^{ \pm} \rightarrow K^{ \pm} \pi^{+} \pi^{-} J / \psi \text { decays }
$$

$$
\begin{array}{r}
\cdot M=3872.0 \pm 0.6 \pm 0.5 \mathrm{MeV} \quad \Gamma<2.3 \mathrm{MeV} \text { at } 90 \% \text { C.L. } \\
\text { width consistent with detector resolution. }
\end{array}
$$

> 1. $D^{0} D^{* 0}$ molecule
> 2. A charmonium hybrid
> 3. $1^{3} D_{2}$ state?

## DOD*O molecule

| Quantity | MeV | $\mathrm{M}_{\mathrm{X}}-\mathrm{M}_{\text {threshold }}$ |
| :--- | :--- | :--- |
| $\mathrm{M}_{\mathrm{X}}$ | $3871.8 \pm 0.7 \pm 0.4$ |  |
| $\mathrm{M}_{\mathrm{D}^{0}}+\mathrm{M}_{\mathrm{D}^{* 0}}$ | $3871.5 \pm 0.7$ | $+0.3 \pm 1.1$ |
| $\mathrm{M}_{\mathrm{D}^{+}}+\mathrm{M}_{\mathrm{D}^{++}}$ | $3879.5 \pm 0.7$ | $-7.7 \pm 1.1$ |

. The mass of the state is right at the $D^{0} D^{* 0}$ threshold! This suggests a loosely bound $D^{0} D^{* 0}$ molecule, right below the dissociation energy
"Molecular Charmonium" discussed in literature since 1975

## Charmonium Options for the $X(3872)$

- Consider all 1D and 2P cc possibilities
- Assume M=3872 MeV
- calculate radiative widths and -strong decay widths


## Strong Decays:

1. Zweig-allowed open-charm decays (DD)
expect $1^{3} D_{2}$ and $1^{1} D_{2}$ but $1^{3} D_{3}$ also narrow because of angular momentum barrier
2. Annihilation type decays
summarized in Ref.[50]. Expressions for decay widths relevant to the 1D and $2 \mathrm{P} c \bar{c}$ states in particular are:

$$
\begin{align*}
& \Gamma\left({ }^{3} \mathrm{D}_{\mathrm{J}} \rightarrow \operatorname{ggg}\right)=\frac{10 \alpha_{s}^{3}}{9 \pi} C_{J} \frac{\left|R_{\mathrm{D}}^{\prime \prime}(0)\right|^{2}}{m_{Q}^{6}} \ln \left(4 m_{Q}\langle r\rangle\right)(7) \\
& \Gamma\left({ }^{1} \mathrm{D}_{2} \rightarrow \mathrm{gg}\right)=\frac{2 \alpha_{s}^{2}}{3} \frac{\left|R_{\mathrm{D}}^{\prime \prime}(0)\right|^{2}}{m_{Q}^{6}}  \tag{8}\\
& \Gamma\left({ }^{3} \mathrm{P}_{2} \rightarrow \mathrm{gg}\right)=\frac{8 \alpha_{s}^{2}}{5} \frac{\left|R_{\mathrm{P}}^{\prime}(0)\right|^{2}}{m_{Q}^{4}} \tag{9}
\end{align*}
$$

3. Hadronic transitions

## Radiative transitions:

$$
\begin{array}{r}
\left.\Gamma\left(\mathrm{n}^{2 \mathrm{~S}+1} \mathrm{~L}_{\mathrm{J}} \rightarrow \mathrm{n}^{\prime 2 \mathrm{~S}^{\prime}+1} \mathrm{~L}_{\mathrm{J}^{\prime}}^{\prime}+\gamma\right)=\frac{4}{3} e_{c}^{2} \alpha \omega^{3} C_{f i} \delta_{\mathrm{SS}^{\prime}}\left|\left\langle\mathrm{n}^{\prime 2 \mathrm{~S}^{\prime}+1} \mathrm{~L}_{\mathrm{J}^{\prime}}^{\prime}\right| r\right| \mathrm{n}^{2 \mathrm{~S}+1} \mathrm{LJ}^{\prime}\right\rangle\left.\right|^{2}, \\
C_{f i}=\max \left(\mathrm{L}, \mathrm{~L}^{\prime}\right)\left(2 \mathrm{~J}^{\prime}+1\right)\left\{\begin{array}{c}
\mathrm{L}^{\prime} \mathrm{J}^{\prime} \mathrm{S} \\
\mathrm{~J} \\
\mathrm{~L}
\end{array}\right\}^{2} .
\end{array}
$$

TABLE II: Radiative transitions in scenario 1: Predictions for the E1 transitions $1 \mathrm{D} \rightarrow 1 \mathrm{P}, 2 \mathrm{P} \rightarrow 2 \mathrm{~S}, 2 \mathrm{P} \rightarrow 1 \mathrm{~S}$ and $2 \mathrm{P} \rightarrow 1 \mathrm{D}$, assuming in all cases that the initial $c \bar{c}$ state has a mass of 3872 MeV . The matrix elements were obtained using the wavefunctions of the Godfrey-Isgur model, Ref.[17]. Unless otherwise stated, the widths are given in keV and the final $c \bar{c}$ masses are PDG values [38].

| Initial <br> state X(3872) | Final state | $\begin{gathered} M_{f} \\ (\mathrm{MeV}) \end{gathered}$ | (MeV) | $\begin{gathered} \langle f\| r\|i\rangle \\ \left(\mathrm{GeV}^{-1}\right) \\ \hline \end{gathered}$ | $C_{f i}$ | Width <br> (keV) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $1^{3} \mathrm{D}_{3}$ | $\chi_{c 2}\left(1^{3} \mathrm{P}_{2}\right) \gamma$ | 3556.2 | 303 | 2.762 | $\frac{2}{5}$ | 367 |
| $1^{3} \mathrm{D}_{2}$ | $\begin{aligned} & \chi_{c 2}\left(1^{3} \mathrm{P}_{2}\right) \gamma \\ & \chi_{c 1}\left(1^{3} \mathrm{P}_{1}\right) \gamma \end{aligned}$ | $\begin{aligned} & 3556.2 \\ & 3510.5 \end{aligned}$ | $\begin{aligned} & 303 \\ & 345 \end{aligned}$ | $\begin{aligned} & 2.769 \\ & 2.588 \end{aligned}$ | $\begin{aligned} & \frac{1}{10} \\ & \frac{3}{10} \end{aligned}$ | 92 356 |
| $1^{3} \mathrm{D}_{1}$ | $\begin{aligned} & \chi_{c 2}\left(1^{3} \mathrm{P}_{2}\right) \gamma \\ & \chi_{c 1}\left(1^{3} \mathrm{P}_{1}\right) \gamma \\ & \chi_{c 0}\left(1^{3} \mathrm{P}_{0}\right) \gamma \end{aligned}$ | $\begin{aligned} & 3556.2 \\ & 3510.5 \\ & 3415 \end{aligned}$ | $\begin{aligned} & 303 \\ & 345 \\ & 430 \end{aligned}$ | $\begin{aligned} & 2.769 \\ & 2.598 \\ & 2.390 \end{aligned}$ | $\begin{aligned} & \frac{1}{90} \\ & \frac{1}{6} \\ & \frac{2}{9} \end{aligned}$ | $\begin{gathered} 10.2 \\ 199 \\ 437 \end{gathered}$ |
| $1{ }^{1} \mathrm{D}_{2}$ | $h_{c}\left(1{ }^{1} \mathrm{P}_{1}\right) \gamma$ | $3517^{\text {a }}$ | 339 | 2.627 | $\frac{2}{5}$ | 464 |

TABLE IV: Partial widths and branching fractions for strong and electromagnetic transitions in scenario 1: We assume in all cases that the initial $c \bar{c}$ state has a mass of 3872 MeV . Details of the calculations are given in the text.

| Initial state | Final state | $\begin{aligned} & \text { Width } \\ & (\mathrm{MeV}) \\ & \hline \end{aligned}$ | B.F. <br> (\%) |
| :---: | :---: | :---: | :---: |
| $1^{3} \mathrm{D}_{3}$ | DD | 4.04 | 84.2 |
|  | ggg | 0.18 | 3.8 |
|  | $J / \psi \pi \pi$ | $0.21 \pm 0.11$ | 4.4 |
|  | $\chi_{0.0}\left(11^{3} \mathrm{P}_{5}\right) \gamma$ | 0.37 | 7.7 |
|  | Total | 4.80 | 100 |
| $1^{3} \mathrm{D}_{2}$ | ggg | 0.08 | 10.8 |
|  | $J / \psi \pi \pi$ | $0.21 \pm 0.11$ | 28.4 |
|  | $\chi_{c 2}\left(1^{3} \mathrm{P}_{2}\right) \gamma$ | 0.09 | 12.2 |
|  | $\chi_{c 1}\left(1^{3} \mathrm{P}_{1}\right) \gamma$ | 0.36 | 48.6 |
|  | Total | 0.74 | 100 |
| $1^{3} \mathrm{D}_{1}$ | DD | 184 | 98.9 |
|  | ggg | 1.15 | 0.6 |
| too wide | $J / \psi \pi \pi$ | $021 \pm 0.11$ | 0.1 |
|  | $\chi_{0}\left(11^{3} P_{1}\right) \gamma$ | 0.20 | 0.1 |
|  | $\chi_{c o}\left(1^{3} \mathrm{P}_{0}\right) \gamma$ | 0.44 | 0.2 |
|  | Total | 186 | 100 |
| $1^{1} \mathrm{D}_{2}$ | $g g$ | 0.19 | 22.1 |
|  | $\eta_{c} \pi \pi$ | $0.21 \pm 0.11$ | 24.4 |
|  | $h_{c}\left(1^{1} \mathrm{P}_{1}\right) \gamma$ | 0.46 | 53.5 |
|  | Total | 0.86 | 100 |


| $2^{3} \mathrm{P}_{2}$ | DD | 21.1 | 80.4 |
| :---: | :---: | :---: | :---: |
|  | $g g$ | 4.4 | 17.2 |
| too wid | $\psi^{\prime}\left(2^{3} \mathrm{~S}_{1}\right) \gamma$ | 8.06 | 0.2 |
|  | $4 ¢\left(1^{5} S_{1}\right)$ | 0.04 | 0.2 |
|  | Total | 25.6 | 100 |
| $2^{3} \mathrm{P}_{1}$ | $q \bar{q} g$ | 1.65 | 95.9 |
|  | $\psi^{\prime}\left(2^{3} \mathrm{~S}_{1}\right) \gamma$ | 0.06 | 3.5 |
|  | $J / \psi\left(1^{3} S_{1}\right)$ | 0.01 | 0.6 |
|  | Total | 1.72 | 100 |
| $2^{3} \mathrm{P}_{0}$ | DD | 13.7 | 24.6 |
|  | $g g$ | 42. | 75.3 |
| too wid | $\psi^{\prime}\left(2^{3} \mathrm{~S}_{1}\right) \gamma$ | 0.07 | $0.1$ |
|  | \% $\left(1^{3} \mathrm{D}_{1}\right.$ | 0.02 | $4 \times 10^{-2}$ |
|  | Total | 55.8 | 100 |
| $2^{1} \mathrm{P}_{1}$ | ggg | 1.29 | 81.6 |
|  | $g g \gamma$ | 0.13 | 8.2 |
|  | $\eta_{c}^{\prime}\left(2^{1} \mathrm{~S}_{0}\right) \gamma$ | 0.09 | 5.7 |
|  | $\eta_{c}\left(1^{1} \mathrm{~S}_{0}\right) \gamma$ | 0.07 | 4.4 |
|  | Total | 1.58 | 100 |

$1^{3} D_{2}$ and $1^{1} D_{2}$ and $1^{3} D_{3}$

| $1^{3} \mathrm{D}_{2}$ | $g g g$ | 0.08 | 10.8 |
| :--- | :--- | :--- | ---: |
|  | $J / \psi \pi \pi$ | $0.21 \pm 0.11$ | 28.4 |
|  | $\chi_{c 2}\left(1^{3} \mathrm{P}_{2}\right) \gamma$ | 0.09 | 12.2 |
|  | $\chi_{c 1}\left(1^{3} \mathrm{P}_{1}\right) \gamma$ | 0.36 | 48.6 |
|  | Total | 0.74 | 100 |
|  |  |  |  |
|  | $\ldots g \ldots$ | $\ldots \sim$ | $\ldots v$ |
| $1^{1} \mathrm{D}_{2}$ | $g g$ | 0.19 | 22.1 |
|  | $\eta_{c} \pi \pi$ | $0.21 \pm 0.11$ | 24.4 |
|  | $h_{c}\left(1^{1} \mathrm{P}_{1}\right) \gamma$ | 0.46 | 53.5 |
|  | Total | 0.86 | 100 |


|  |  |  |  |
| :--- | :--- | :--- | ---: |
| $1^{3} \mathrm{D}_{3}$ | DD | 4.04 | 84.2 |
|  | $g g g$ | 0.18 | 3.8 |
|  | $J / \psi \pi \pi$ | $0.21 \pm 0.11$ | 4.4 |
|  | $\chi_{c 2}\left(1^{3} \mathrm{P}_{2}\right) \gamma$ | 0.37 | 7.7 |
|  | lotal | 4.80 | 100 |

$2^{3} P_{1}$ and $2^{1} P_{1}$

|  |  | $\cdots$ | $\cdots$ |
| :--- | :--- | ---: | ---: |
| $2^{3} \mathrm{P}_{1}$ | $q \bar{q} g$ | 1.65 | 95.9 |
|  | $\psi^{\prime}\left(2^{3} \mathrm{~S}_{1}\right) \gamma$ | 0.06 | 3.5 |
|  | $J / \psi\left(1^{3} S_{1}\right) \gamma$ | 0.01 | 0.6 |
|  | Total | 1.72 | 100 |
|  | $\cdots g g$ | $\cdots \cdots$ | $\cdots$ |
| $2^{1} \mathrm{P}_{1}$ | $g g g$ | 1.29 | 81.6 |
|  | $g g \gamma$ | 0.13 | 8.2 |
|  | $\eta_{c}^{\prime}\left(2^{1} \mathrm{~S}_{0}\right) \gamma$ | 0.09 | 5.7 |
|  | $\eta_{c}\left(1^{1} \mathrm{~S}_{0}\right) \gamma$ | 0.07 | 4.4 |
|  | Total | 1.58 | 100 |

The problem here is that the $B R$ to $\gamma$ and $\pi \pi$ is quite small and not the final states being looked for

Consider the charmonium possibilities:
T.Barnes,S.Godfrey, PR D69, 050400 (2004)

Eichten, Lane, Quigg, PR D69, 094019 (2004) Barnes, Godfrey, Swanson, PR D 054026 (2005)
1D and 2P multiplets only states nearby in mass $1^{1} D_{2} 2^{3} P_{0} 2^{3} P_{1} 2^{3} P_{2}$ have $C=+$

But $\mathrm{X}(3872) \rightarrow \gamma \mathrm{J} / \psi$ implies $C=+$ Belle [hep-ex/0505037] Babar Gowdy Moriond talk
Angular distributions favour JPC $=1^{++}$Belle [hep-ex/0505038]
The unique surviving charmonium candidate is $2^{3} \mathrm{P}_{1}$
BUT identification of $Z(3931)$ with $2^{3} \mathrm{P}_{2}$ implies 2P mass ~ 3940 MeV

## $D^{0} D^{* 0}$ molecule or "tetraquark" is a popular/likely explanation: see Voloshin

## $y(4260)$

Discovered by Babar as enhancement in $\pi \pi \mathrm{J} / \psi$ subsystem in $\mathrm{e}^{+} \mathrm{e}^{-} \rightarrow \gamma_{\text {ISR }} \psi \pi \pi$ PRL 95, 142001(2005) 「hen-ex/05060811
$M=4259 \pm 8 \pm 4 \mathrm{MeV}$
$\Gamma=88 \pm 23 \pm 5 \mathrm{MeV}$
$\Gamma_{\text {ee }} \times \mathrm{BR}\left(\mathrm{Y} \rightarrow \pi^{+} \pi^{-} \mathrm{J} / \psi\right)=5.5 \pm 1.0 \pm 0.8 \mathrm{eV}$
ISR production tells us JPC $=1-$
Further evidence in
$\mathrm{B} \rightarrow \mathrm{K}\left(\pi^{+} \pi^{-} \mathrm{J} / \psi\right)$ PR D73, 011101(2006)


Confirmed by CLEO
hep-ex/0602034


-The first unaccounted 1-- state is the $\psi(3 D)$
-Quark models estimate $M(\psi(3 D)) \sim 4500 \mathrm{MeV}$ much too heavy for the $Y(4260)$
$Y(4260)$ represents an overpopulation of expected 1-- states

Absence of open charm production also against conventional cc state

Other explanations are:

- $4(45) \quad$ Phys Rev D72, 031503 (2005)
- Tetraquark Phys Rev D72, 031502 (2005)
-CC hybrid Phys Lett B625, 212 (2005);
Phys Lett B628, 215 (2005)
Phys Lett B631, 164 (2005)


## Y(4260): Hybrid?

-Flux tube model predicts lowest cc hybrid at 4200 MeV
-LGT expects lowest cc hybrid at 4200 MeV [Phys Lett B401, 308 (1997)]

- Models of hybrids say $\Psi(0)=0$ so would have small $e^{+} e^{-}$width
-LGT found bb hybrids have large couplings to closed flavour modes
- Similar to BaBar observation of $\mathrm{Y} \rightarrow \pi^{+} \pi^{-\mathrm{J} /} / \mathrm{F}$ :

$$
\begin{aligned}
& \mathrm{BR}\left(\mathrm{Y} \rightarrow \pi^{+} \pi^{-\mathrm{J} / \psi)}\right)>8.8 \% \\
& \Gamma\left(\mathrm{Y} \rightarrow \pi^{+} \pi^{-\mathrm{J}} / \psi\right)>7.7 \pm 2.1 \mathrm{MeV}
\end{aligned}
$$

- Much larger than typical charmonium transitions:

$$
\Gamma\left(\psi(3770) \rightarrow \pi^{+} \pi^{-J} / \psi\right) \sim 80 \mathrm{keV}
$$

$\cdot \mathrm{Y}$ is seen while $\psi(4040), \psi(4160) \psi(4415)$ are no $\dagger$

## How to test $\mathrm{Y}(4260)$ hybrid assignment:

## Decays:

-LGT study suggest searching for other closed charm modes with

$$
{ }^{\mathrm{JPC}=1-} J / \psi \eta, J / \psi \eta^{\prime}, \chi_{J} \omega \ldots
$$

- Models predict the dominant hybrid charmonium open-charm decay modes will be a meson pair with

S-wave ( $D, D^{*}, D_{s}, D_{s}{ }^{*}$ ) + P-wave $\left(D_{J}, D_{s J}\right)$
-The dominant decay mode expected to be $D+D_{1}(2430)$
$D_{1}(2420)$ has width $\sim 300 \mathrm{MeV}$ and decays to $D^{*} \pi$
-Suggests search for $Y(4260)$ in $D^{*} \pi$

- Evidence of large $D_{1}(2430)$ signal would be strong evidence for hybrid
- But models of hybrids are untested so to be cautious
-If seen in other modes like $D^{*}, D_{s} D_{s}{ }^{*}$ comparable to $\pi^{+} \pi^{-J} / \psi$ maybe still hybrid but decay model not accurate


## Search for Partner States: (fill in the multiplet)

- Mass ca. 4.0-4.5 GeV, with LGT preferring the higher range.
(e.g.: X.Liao and T.Manke, hep-lat/0210030)
- Confirm that no c¢ states with the same $J^{P C}$ are expected at this mass.
-Identify $J^{P C}$ partners of the hybrid candidate nearby in mass.
-The most convincing evidence:
- partners, especially $J^{P C}$ exotics.
-The f-t model expects:

$$
0^{+-}, 1^{-+}, 2^{+-}, 0^{-+}, 1^{+-}, 2^{-+}, 1^{++}, 1^{--}
$$

## Summary

Many new results, considerable progress!

| $D_{S J}(2317)$ | Most likely $0^{+}(c \bar{S})$ |
| :--- | :--- |
| $D_{S J}(2460)$ | Most likely $1^{+}(c \bar{S})$ |
| $D_{S J}(2632)$ | Needs confirmation |
| $X(3872)$ | Molecule? - see Voloshin |
| $X(3943)$ | $\eta^{\prime \prime}{ }_{c}\left(3^{1} S_{0}\right)$-look for $\gamma \gamma \rightarrow D^{*}$ |
| $Y(3943)$ | $\chi_{c 1}^{\prime}\left(2^{3} P_{1}\right)$-look for DD \& DD* |
| $Z(3930)$ | $\chi_{c 2}^{\prime}\left(2^{3} P_{2}\right)$-confirm by DD* |
| $Y(4260)$ | Hybrid? |

- Much more to learn; ie search for $1^{3} D 31^{3} D 21^{1} D 21^{3} F 21^{3} F 4$

